

The ABC's of Nuclear Science*

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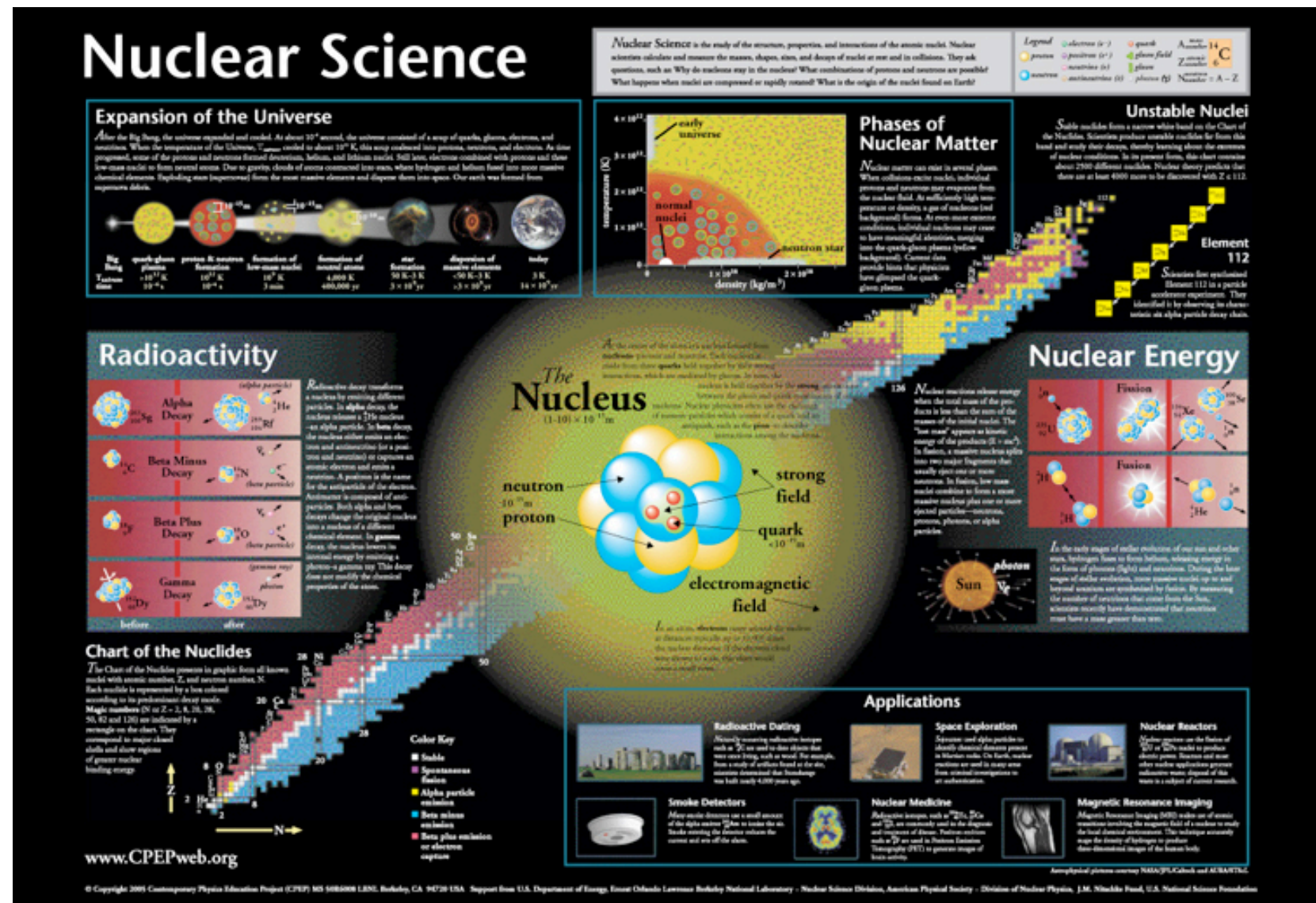
awpoon@lbl.gov

<http://neutrino.lbl.gov>

* Slides adapted from Rick Norman's previous talks in this workshop series

Outline

- A brief history of the nucleus
- A walk around the Nuclear Science wall chart
- Resources



<http://www.lbl.gov/abc>

The classical elements



Earth



Air



Greek

+

Aether
(Quintessence)



Fire

Water



The classical elements



Earth

Air



Greek



+

Aether
(Quintessence)

Buddhism

Hinduism



Fire

Water



The classical elements



Earth



Air



+

Aether
(Quintessence)



Fire



Water

+



Salt



Sulfur



Mercury

Medieval Alchemy

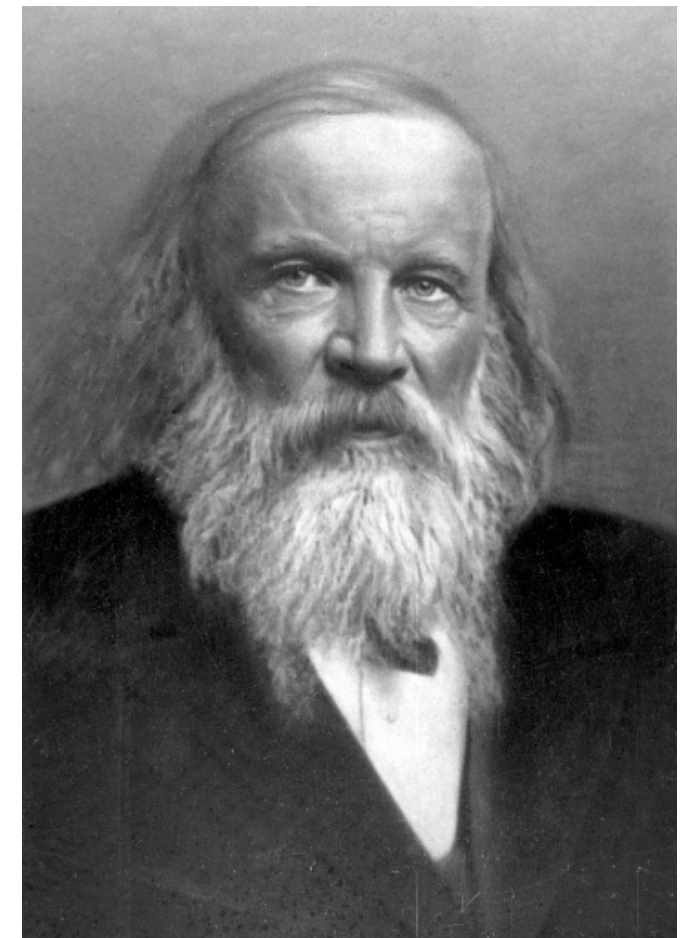
Mendeleev's Periodic Table

ОПЫТЪ СИСТЕМЫ ЭЛЕМЕНТОВЪ.

ОСНОВАННОЙ НА ИХЪ АТОМНОМЪ ВѢСѢ И ХИМИЧЕСКОМЪ СХОДСТВѢ.

			Ti = 50	Zr = 90	? = 180.
			V = 51	Nb = 94	Ta = 182.
			Cr = 52	Mo = 96	W = 186.
			Mn = 55	Rh = 104,4	Pt = 197,1.
			Fe = 56	Ru = 104,4	Ir = 198.
			Ni = Co = 59	Pd = 106,8	Os = 199.
			Cu = 63,4	Ag = 108	Hg = 200.
H = 1	Be = 9,4	Mg = 24	Zn = 65,2	Cd = 112	
	B = 11	Al = 27,1	? = 68	Ur = 116	Au = 197?
	C = 12	Si = 28	? = 70	Sn = 118	
	N = 14	P = 31	As = 75	Sb = 122	Bi = 210?
	O = 16	S = 32	Se = 79,4	Te = 128?	
	F = 19	Cl = 35,5	Br = 80	I = 127	
Li = 7	Na = 23	K = 39	Rb = 85,4	Cs = 133	Tl = 204.
		Ca = 40	Sr = 87,6	Ba = 137	Pb = 207.
			? = 45	Ce = 92	
		?Er = 56	La = 94		
		?Yt = 60	Di = 95		
		?In = 75,6	Th = 118?		

Д. Менделѣевъ



Dmitri Mendeleev
(1834-1907)

I saw in a dream a table where all the elements fell into place as required. Awakening, I immediately wrote it down on a piece of paper. Only in one place did a correction later appear necessary.

-Dmitri Mendeleev, 1869

Periodic Table of the Elements

11.01

H

Hydrogen

36.94

Li

Lithium

6.94

Be

Beryllium

22.99

Na

Sodium

24.31

Mg

Magnesium

39.10

K

Potassium

40.08

Ca

Calcium

44.96

Sc

Scandium

47.90

Ti

Titanium

50.94

V

Vanadium

51.996

Cr

Chromium

54.94

Mn

Manganese

55.85

Fe

Iron

58.93

Co

Cobalt

58.70

Ni

Nickel

63.55

Cu

Copper

65.37

Zn

Zinc

69.72

Ga

Gallium

72.59

Ge

Germanium

74.92

As

Arsenic

78.96

Se

Selenium

79.90

Br

Bromine

83.80

Kr

Krypton

85.47

Rb

Rubidium

87.62

Sr

Strontium

88.91

Y

Yttrium

91.22

Zr

Zirconium

92.91

Nb

Niobium

95.94

Mo

Molybdenum

(98)

Tc

Technetium

101.07

Ru

Ruthenium

102.91

Rh

Rhodium

106.40

Pd

Palladium

107.87

Ag

Silver

112.41

Cd

Cadmium

114.82

In

Indium

118.69

Sn

Tin

121.75

Sb

Antimony

127.60

Te

Tellurium

126.90

I

Iodine

131.30

Xe

Xenon

132.91

Cs

Cesium

137.33

Ba

Barium

138.91

La

Lanthanum

178.49

Hf

Hafnium

180.95

Ta

Tantalum

183.85

W

Tungsten

186.21

Re

Rhenium

190.20

Os

Osmium

192.22

Ir

Iridium

195.09

Pt

Platinum

196.97

Au

Gold

200.59

Hg

Mercury

204.37

Tl

Thallium

207.19

Pb

Lead

208.98

Bi

Bismuth

(209)

Po

Polonium

(210)

At

Astatine

(222)

Rn

Radon

(223)

Fr

Francium

226.03

Ra

Radium

227.03

Ac

Actinium

(261)

Rf

Rutherfordium

(262)

Ha

Hahnium

(266)

Sg

Seaborgium

(266)

Bh

Bohrium

(265)

Hs

Hassium

(266)

Mt

Meitnerium

(271)

(272)

(277)

(113)

(285)

(115)

(289)

(117)

(293)

1.01

H

Hydrogen

10.81

B

Boron

12.01

C

Carbon

14.01

N

Nitrogen

15.999

O

Oxygen

18.998

F

Fluorine

20.18

Ne

Neon

26.98

Al

Aluminum

28.09

Si

Silicon

30.97

P

Phosphorus

32.06

S

Sulfur

35.45

Cl

Chlorine

39.96

Ar

Argon

14

28.09

Si

Silicon

atomic number

atomic weight

symbol

name

black

solid

blue

liquid

red

gas

white

synthetically prepared

most stable isotope

alkali metals

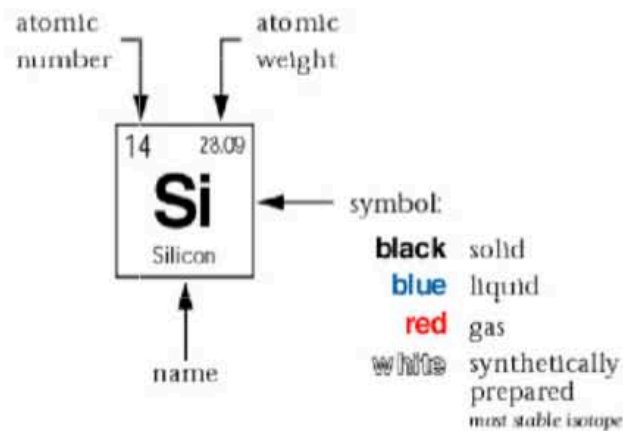
alkaline earth metals

transitional metals

other metals

nonmetals

noble gases



- alkali metals
- alkaline earth metals
- transitional metals
- other metals
- nonmetals
- noble gases

Lanthanide series

58 Ce Cerium	59 Pr Praseodymium	60 Nd Neodymium	61 Pm Promethium	62 Sm Samarium	63 Eu Europium	64 Gd Gadolinium	65 Tb Terbium	66 Dy Dysprosium	67 Ho Holmium	68 Er Erbium	69 Tm Thulium	70 Yb Ytterbium	71 Lu Lutetium
90 Th Thorium	91 Pa Protactinium	92 U Uranium	93 Np Neptunium	94 Pu Plutonium	95 Am Americium	96 Cm Curium	97 Bk Berkelium	98 Cf Californium	99 Es Einsteinium	100 Fm Fermium	101 Md Mendelevium	102 No Nobelium	103 Lr Lawrencium

Actinide series

Viewing the periodic table with x-rays

- Elements are distinguished by the atomic number, Z , which is the number of protons in its nucleus → also determines the number of electrons (charge neutrality).
- Wilhelm Röntgen discovered x-rays from cathode-ray tubes in 1895.
- Charles Glover Barkla discovered that each element has its own characteristic x-ray spectrum.
- Henry Moseley established that Z is related to the frequency of the x-ray:

$$Z \propto \sqrt{f}$$

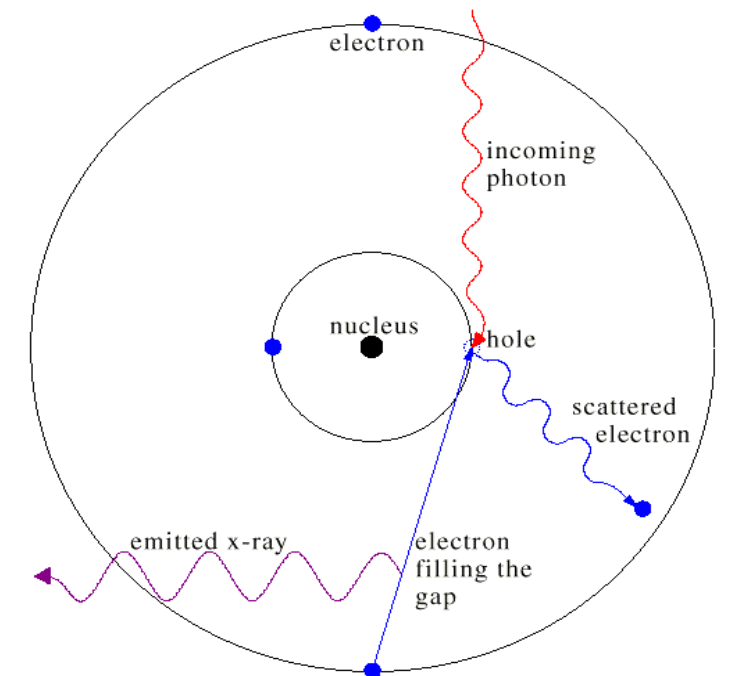


Hand mit Ringen : First “medical” X-ray

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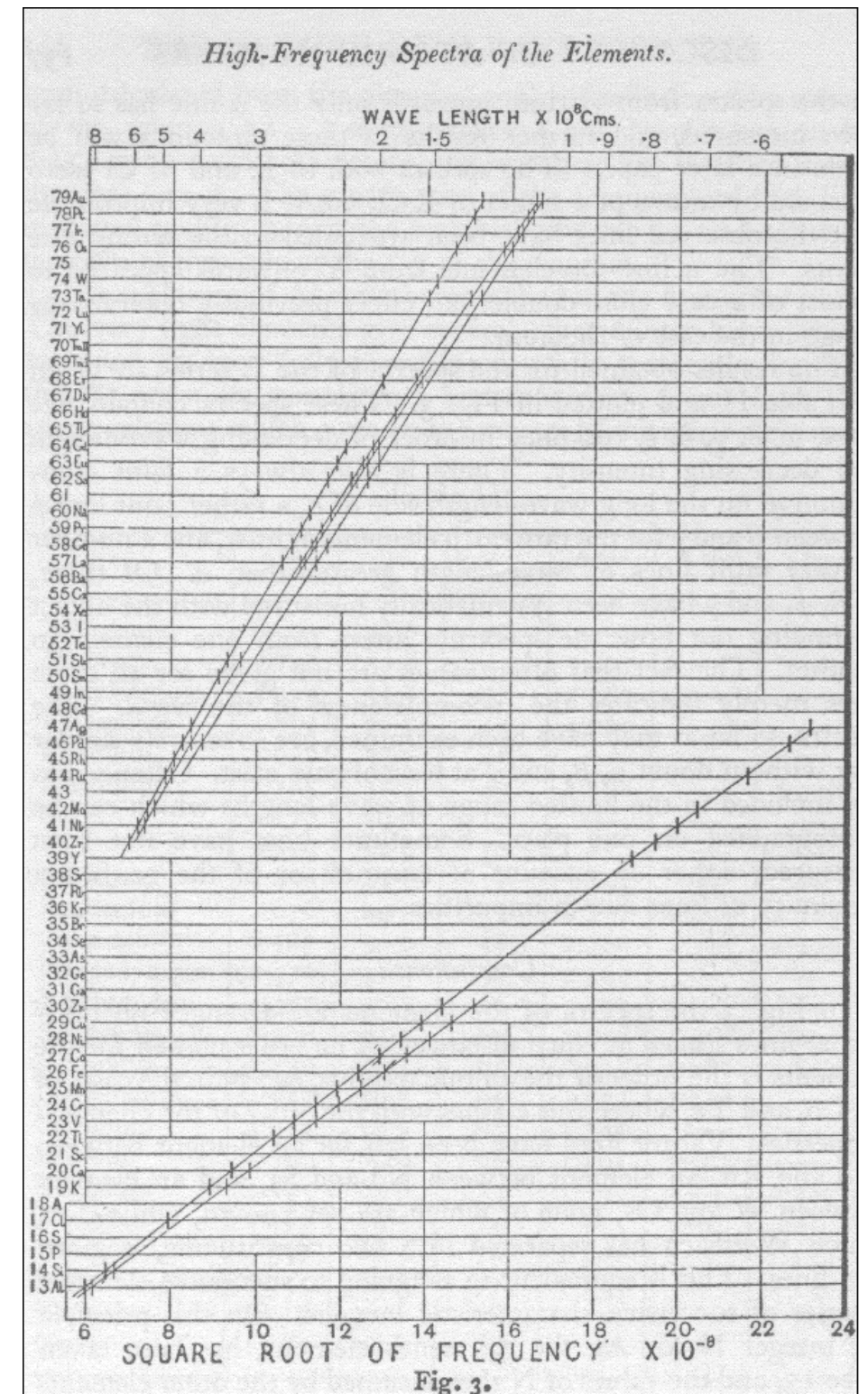


X-ray fluorescence

Viewing the periodic table with x-rays

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A student XRF project at LBNL



<http://ie.lbl.gov/xray>

A student XRF project at LBNL



^{241}Am source
 $E_\gamma = 59.5 \text{ keV}$

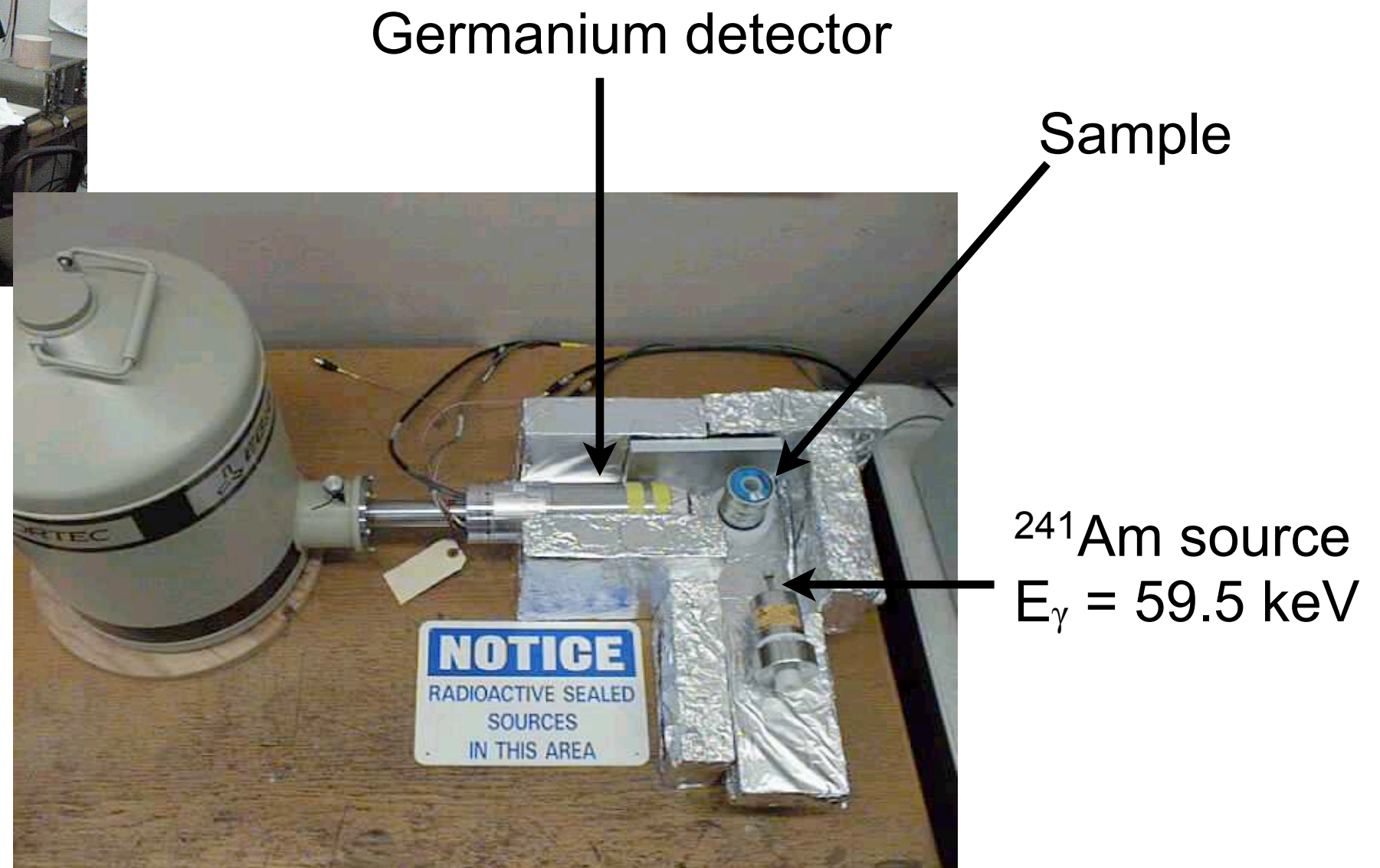
<http://ie.lbl.gov/xray>

A student XRF project at LBNL



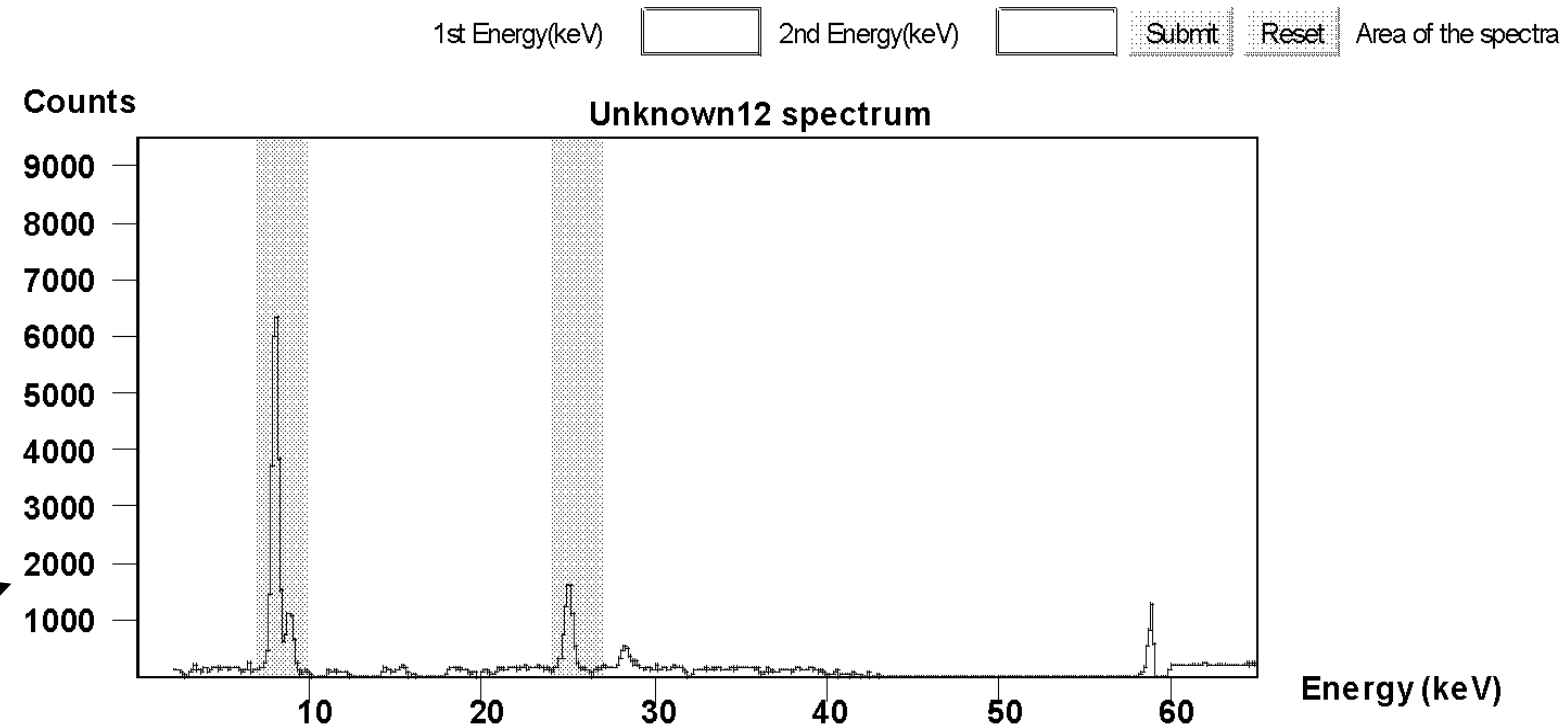
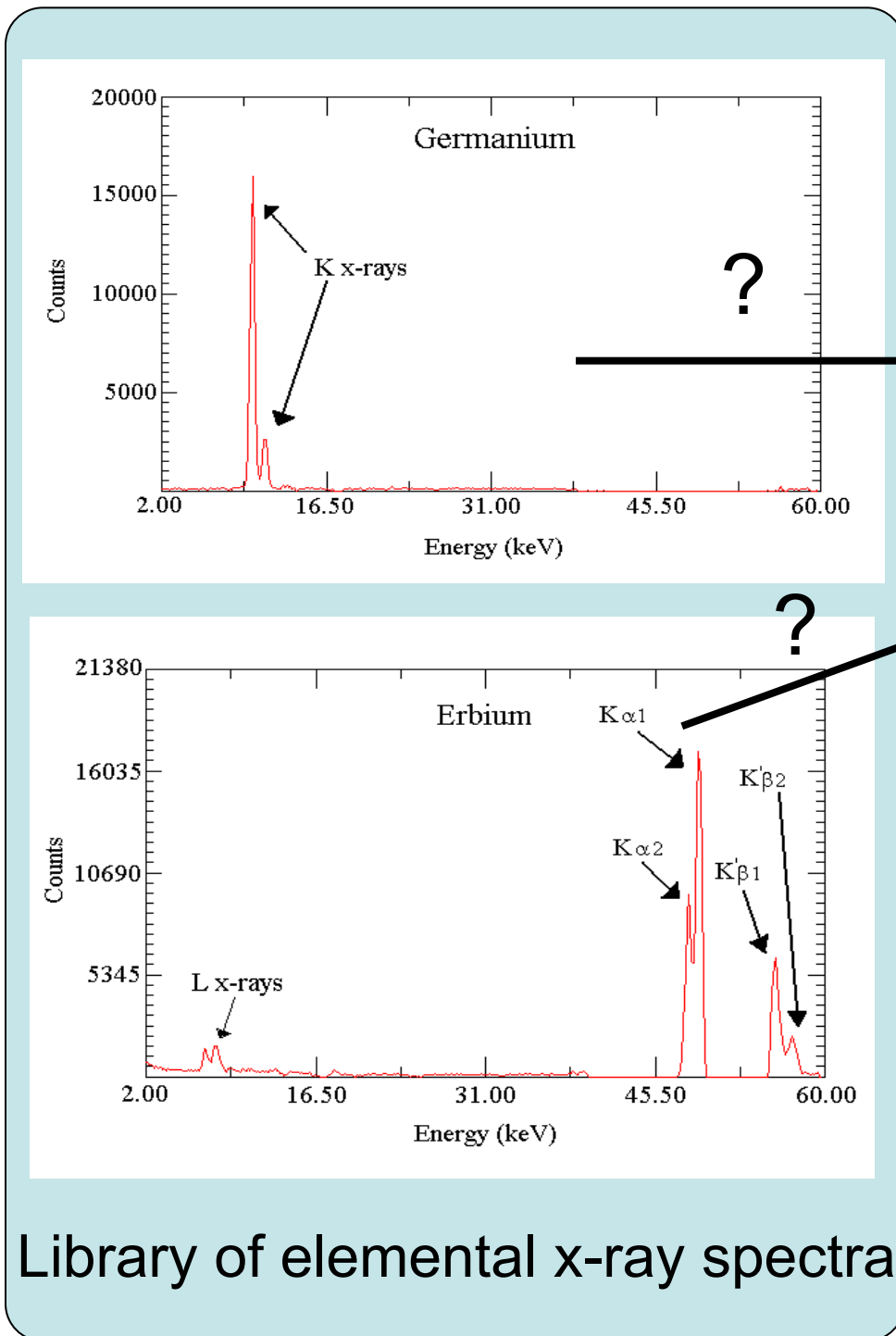
<http://ie.lbl.gov/xray>

A student XRF project at LBNL



<http://ie.lbl.gov/xray>

Examples of x-ray spectra available



- Fluorescence spectrum of the sample can be “unfolded” to find out its elemental make-up.
- XRF is a commonly used technique in many fields

The beginning of Nuclear Science



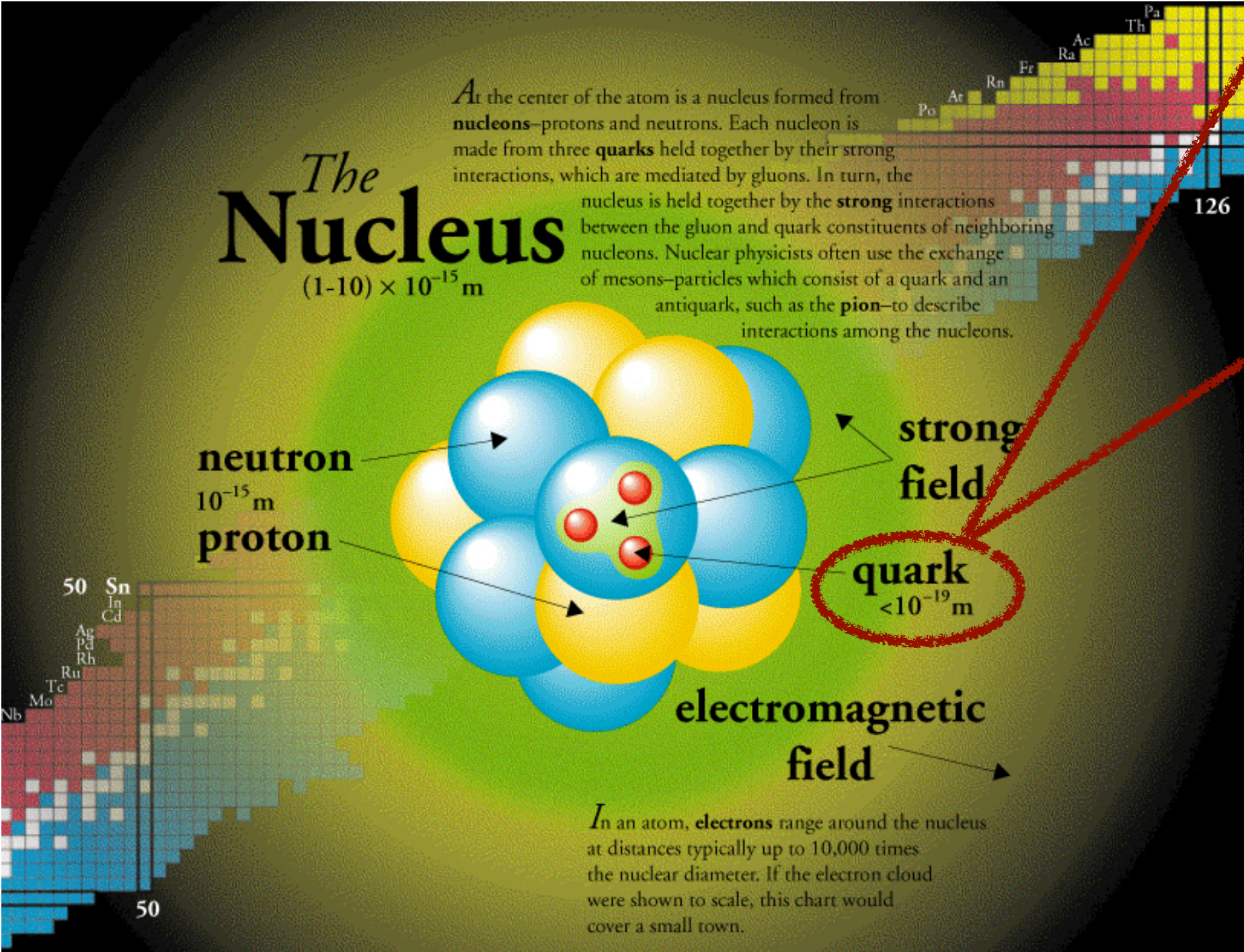
Henri Becquerel



Marie & Pierre Curie

- Scientists in the late 1800's and early 1900's made discoveries which would change the course of science, history and medicine.
- Henri Becquerel: Uranium salt “fogged” a photographic plate. This “Becquerel ray” \neq x-ray. The discovery of radioactivity.
- Marie & Pierre Curie: Extracted uranium from ore; but the left-over had even more radioactivity \rightarrow discovery of polonium and radium

The nucleus



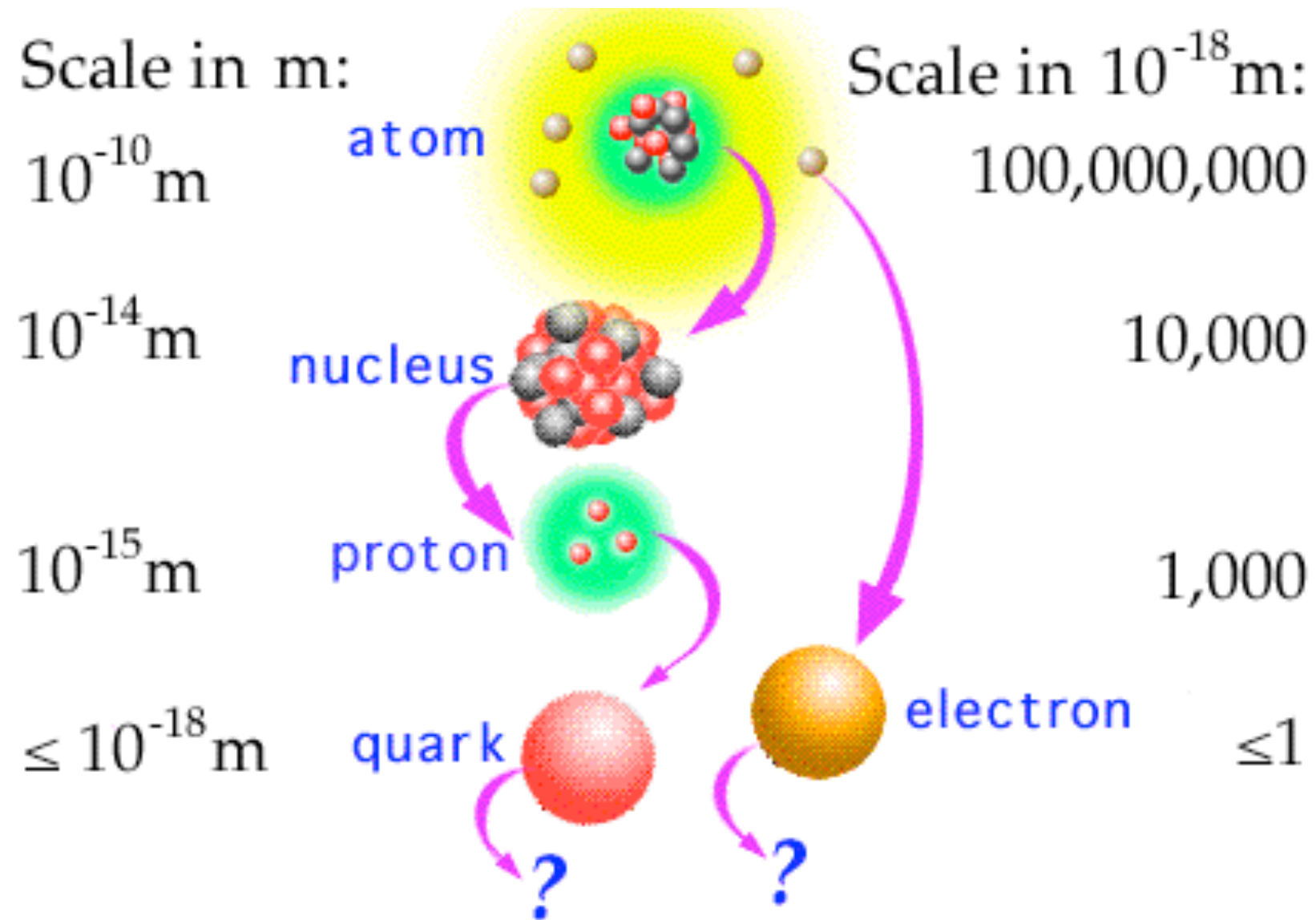
Elementary Particles

Quarks	<i>u</i>	<i>c</i>	<i>t</i>	Force Carriers
	<i>d</i>	<i>s</i>	<i>b</i>	
Leptons	<i>ν_e</i>	<i>ν_μ</i>	<i>ν_τ</i>	<i>Z</i>
	<i>e</i>	<i>μ</i>	<i>τ</i>	<i>W</i>

Three Generations of Matter

proton	u u d
neutron	u d d

The nucleus



Isotopes

- J.J. Thomson sent Ne ions through electric and magnetic fields and saw two “beams” recorded on a photographic plate.
- F.W. Aston perfected the mass spectrograph to discover more isotopes.

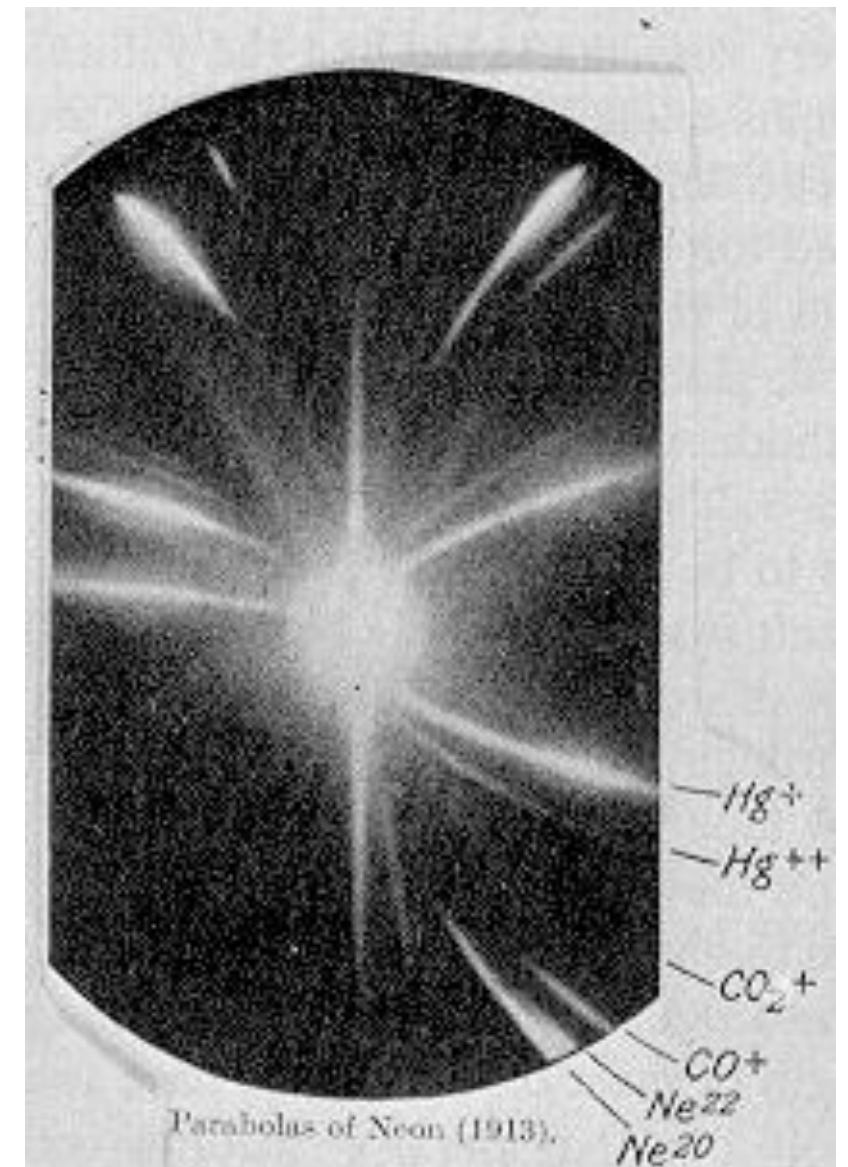


$$A = N + Z$$

A: mass number

N: # of neutrons

Z: atomic number



- Three isotopes of hydrogen:

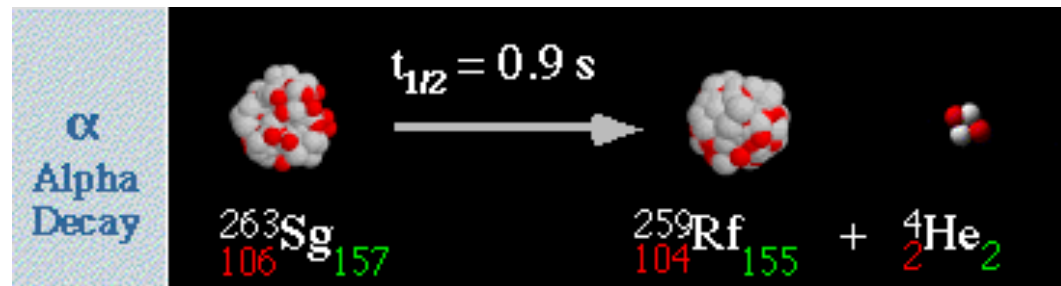
Hydrogen	Deuterium	Tritium
${}^1_1\text{H}$	${}^2_1\text{H}$	${}^3_1\text{H}$

Types of radioactivity

- **Radioactivity:** The spontaneous decay or disintegration of an unstable atomic nucleus accompanied by the emission of radiation.
- Three types:

Types of radioactivity

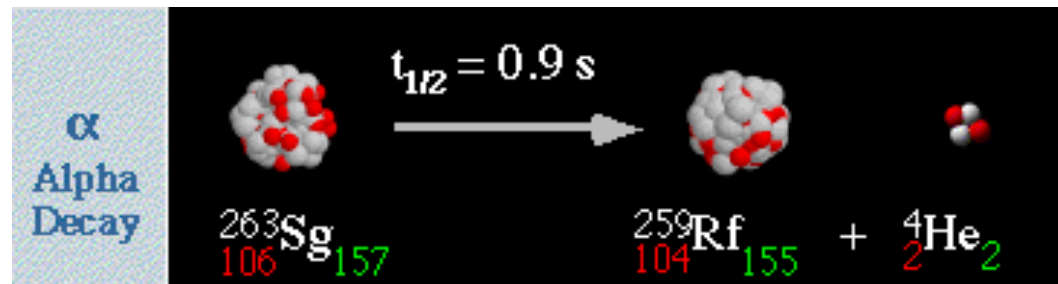
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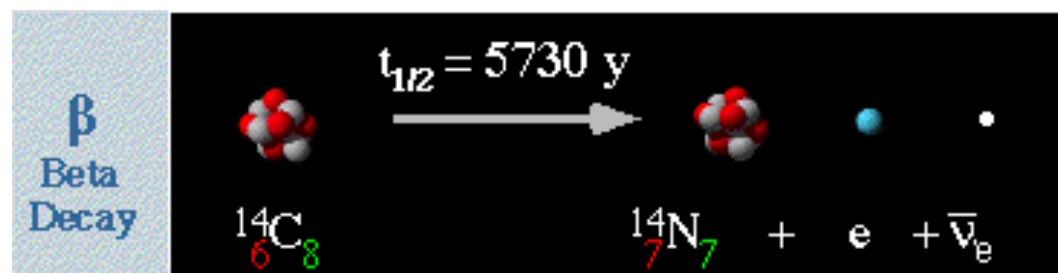
$$\alpha : (A, Z) \rightarrow (A - 4, Z - 2) + {}^4_2\text{He}$$

Types of radioactivity

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$$\alpha : (A, Z) \rightarrow (A - 4, Z - 2) + ^4_2\text{He}$$



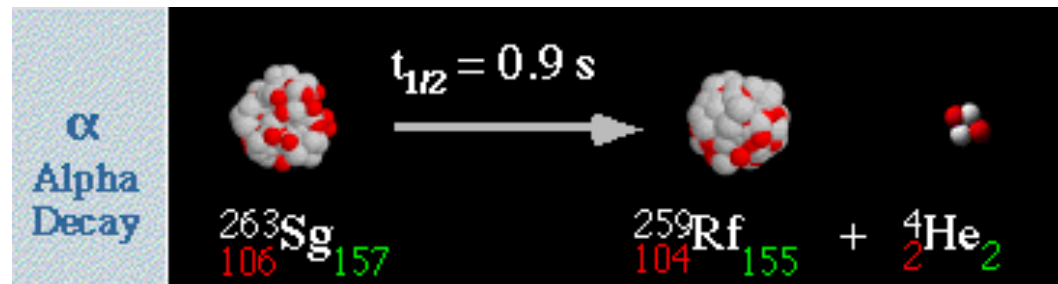
$$\beta^- : (A, Z) \rightarrow (A, Z + 1) + e^- + \bar{\nu}_e$$

$$\beta^+ : (A, Z) \rightarrow (A, Z - 1) + e^+ + \nu_e$$

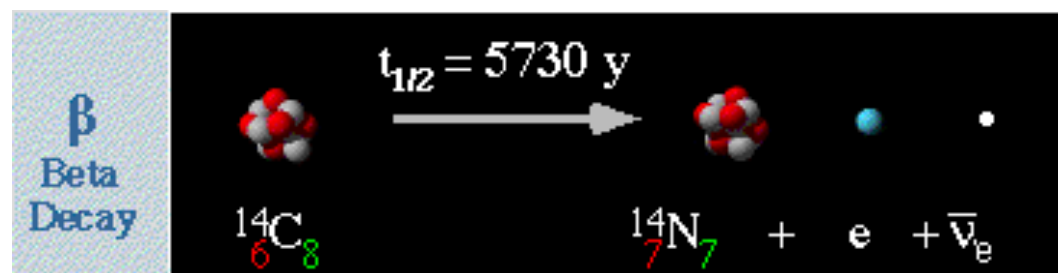
Electron capture : $(A, Z) + e^- \rightarrow (A, Z - 1) + \nu_e$

Types of radioactivity

- **Radioactivity:** The spontaneous decay or disintegration of an unstable atomic nucleus accompanied by the emission of radiation.
- Three types:



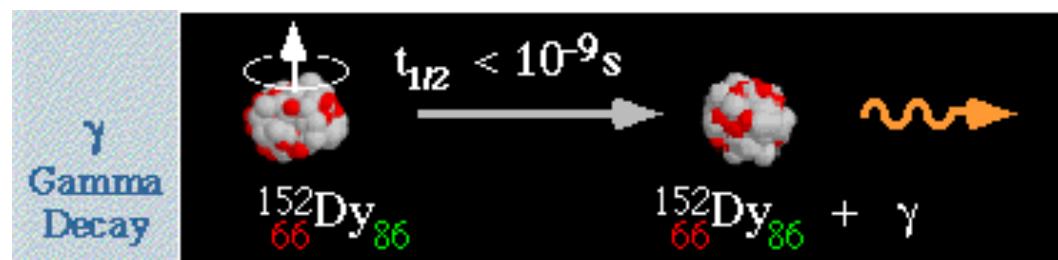
$$\alpha : (A, Z) \rightarrow (A - 4, Z - 2) + ^4_2\text{He}$$



$$\beta^- : (A, Z) \rightarrow (A, Z + 1) + e^- + \bar{\nu}_e$$

$$\beta^+ : (A, Z) \rightarrow (A, Z - 1) + e^+ + \nu_e$$

Electron capture : $(A, Z) + e^- \rightarrow (A, Z - 1) + \nu_e$



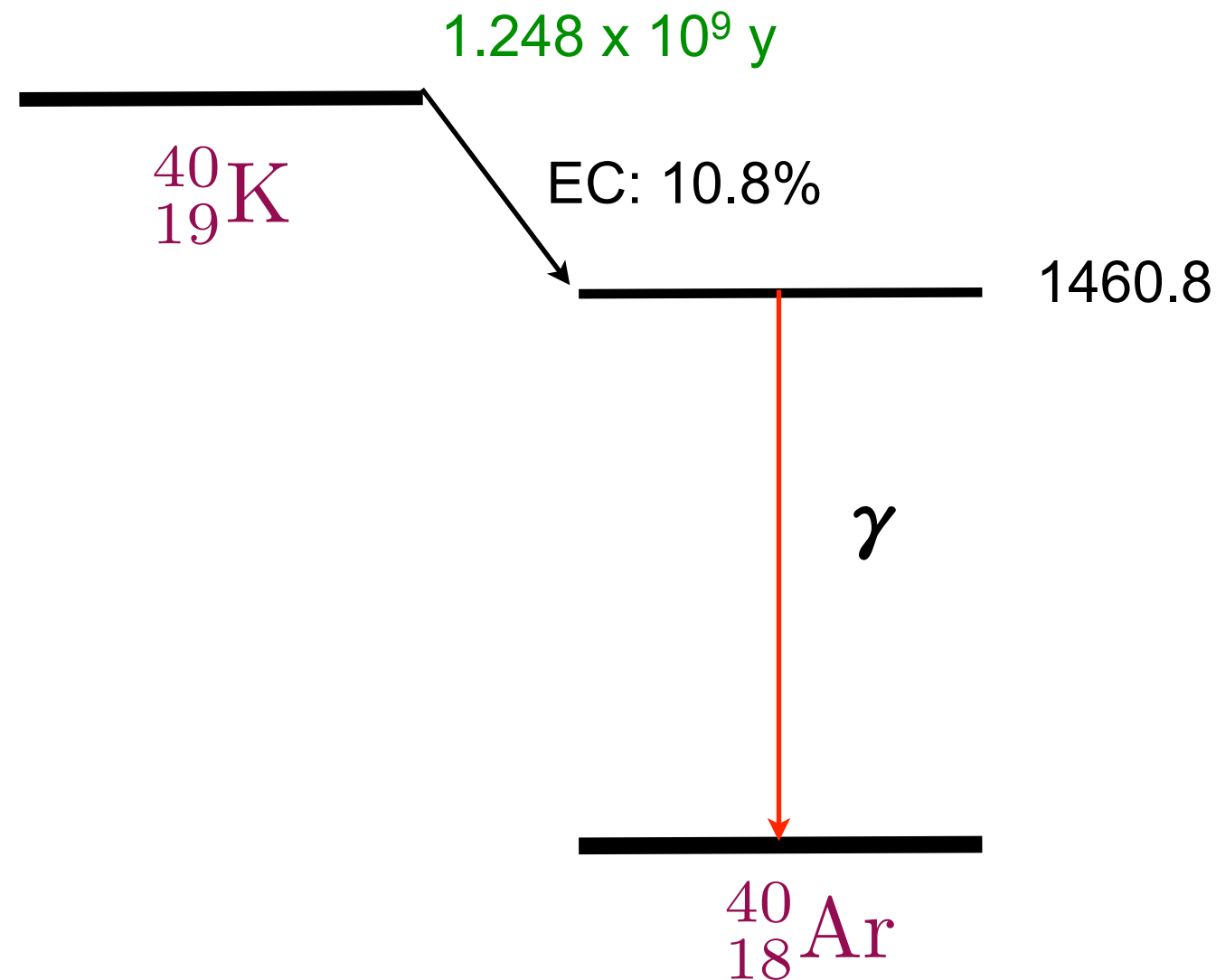
$$\gamma : (A, Z)^* \rightarrow (A, Z) + \gamma$$

- Examples: natural decay chains



Beta and gamma decays

- Example: natural $^{40}_{19}\text{K}$



Natural radioactivity: Do you know that...

- Natural radioactive decays “power” our Earth

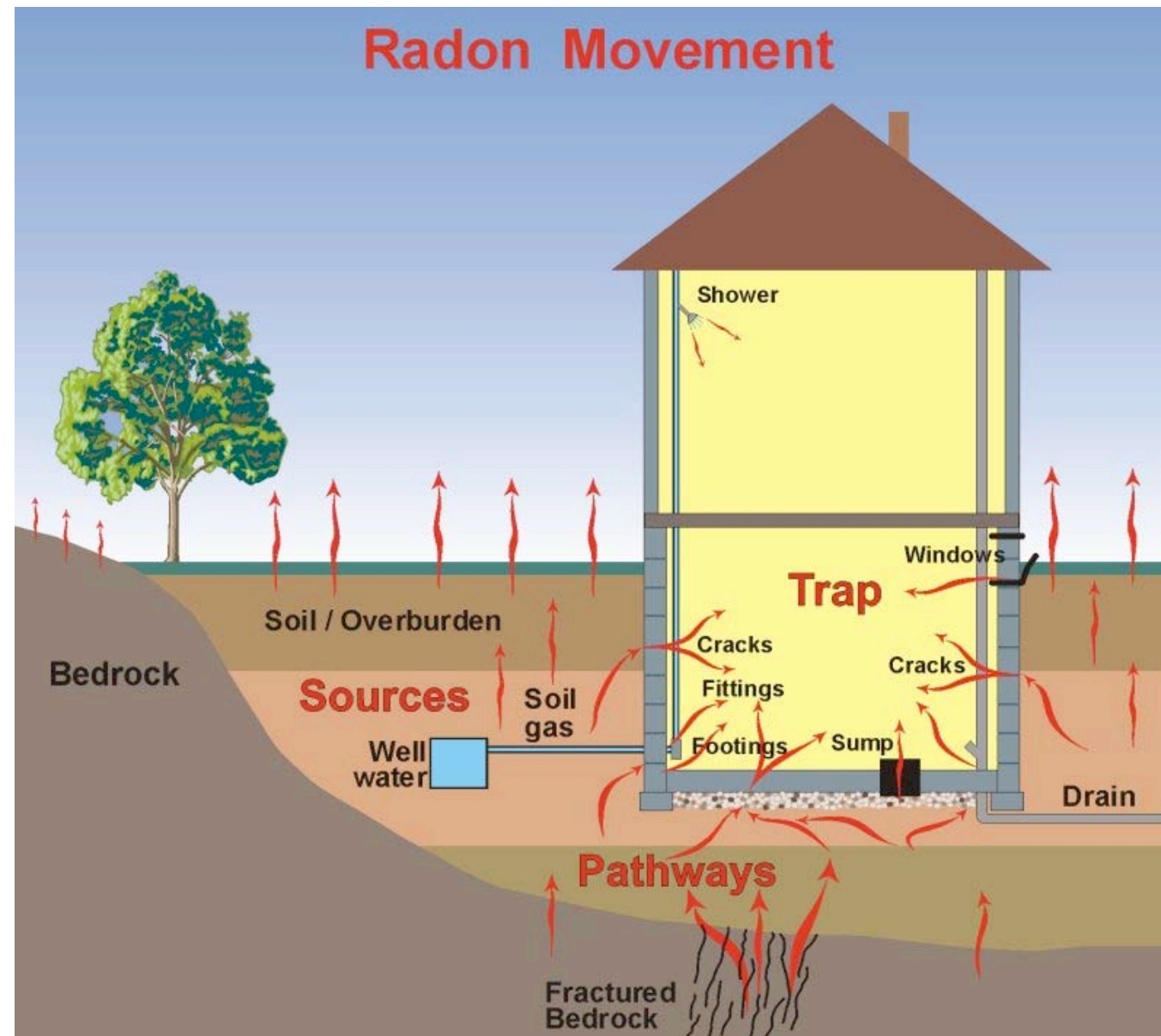
- *How do we know?*

We have observed the neutrinos from the beta decays (“geo-neutrinos”) in the chains.

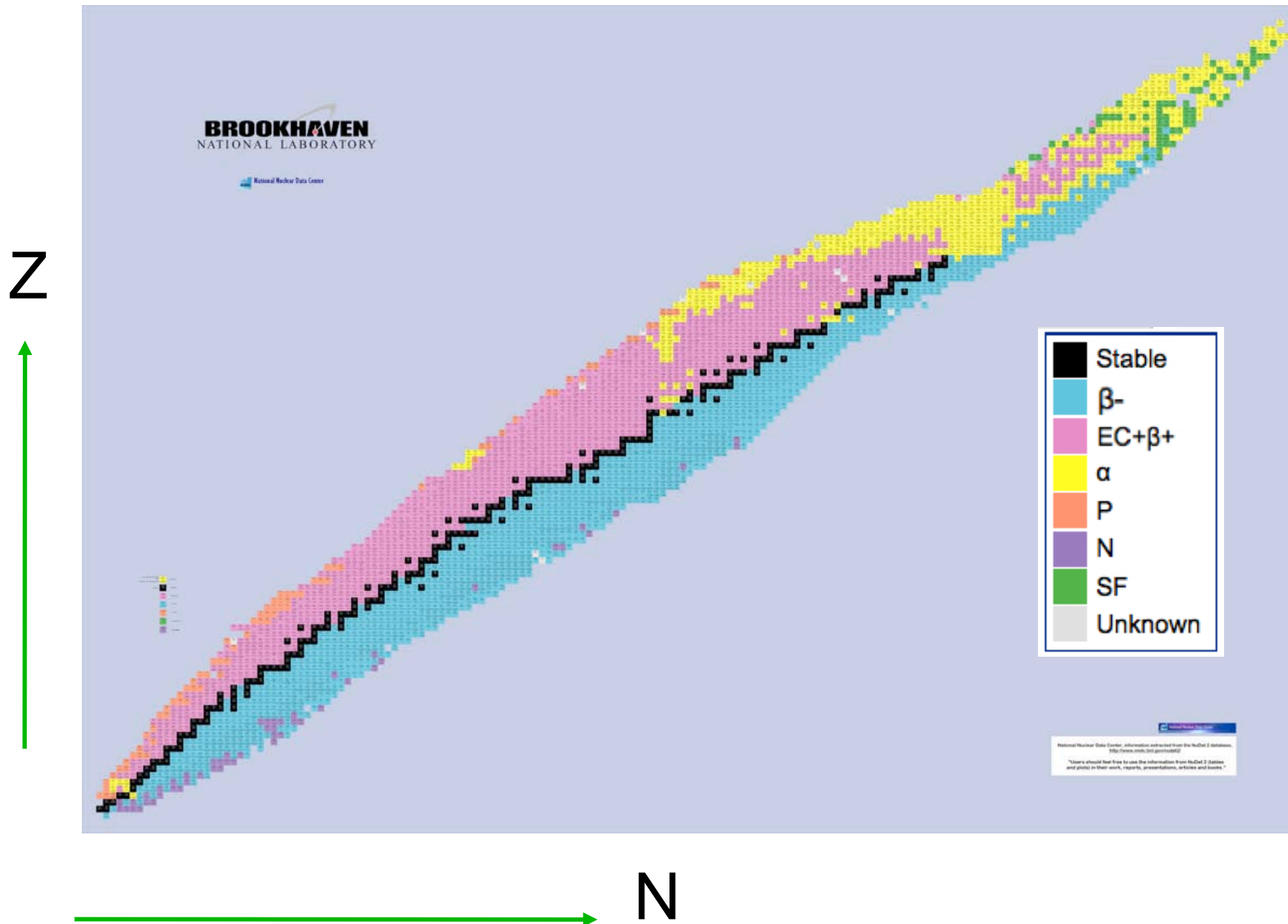


Natural radioactivity: Do you know that...

- Radon is part of the decay chains



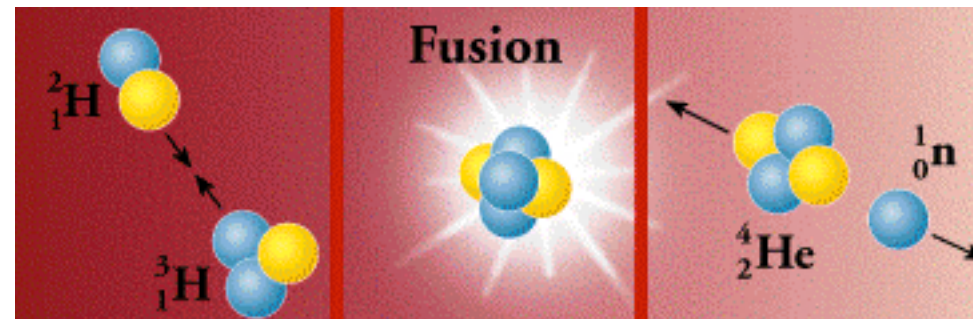
The Chart of the Nuclides: A “2-D” Periodic Table



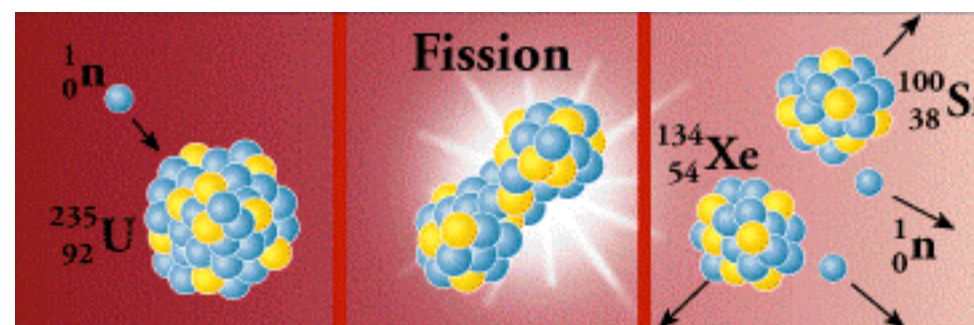
- *Nuclei have shells similar to the electron shells in an atom. The electron shells determine the atomic and chemical properties; the nuclear shells determine the nuclear properties. Nuclei with filled shells (a “magic” number of neutrons or protons) are more stable.*
- *Moving from one box to neighboring boxes can be accomplished through nuclear reactions.*

Nuclear reactions

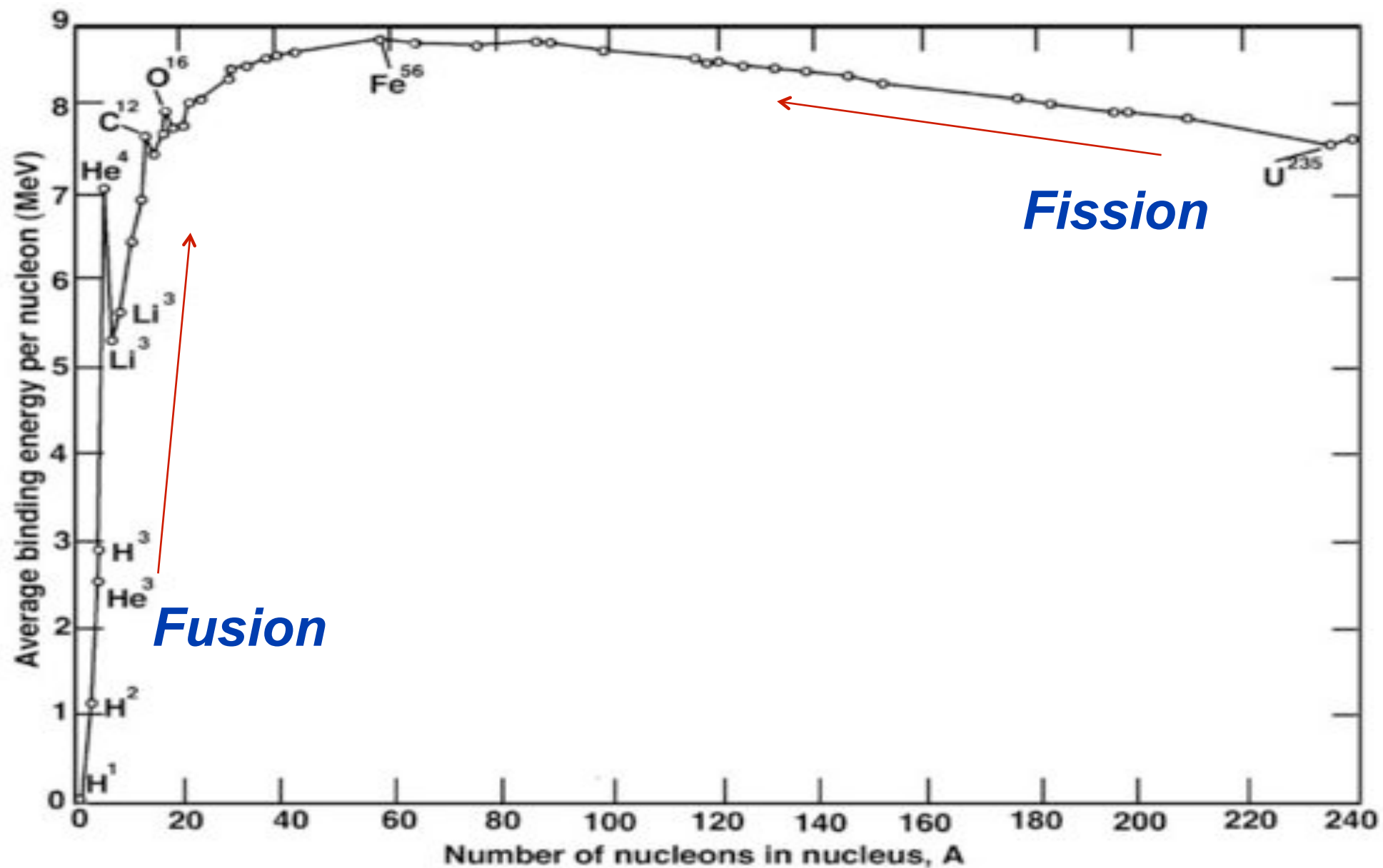
- Examples:
 - **Fusion:** how the Sun generates energy



- **Fission:** how nuclear reactors generate energy

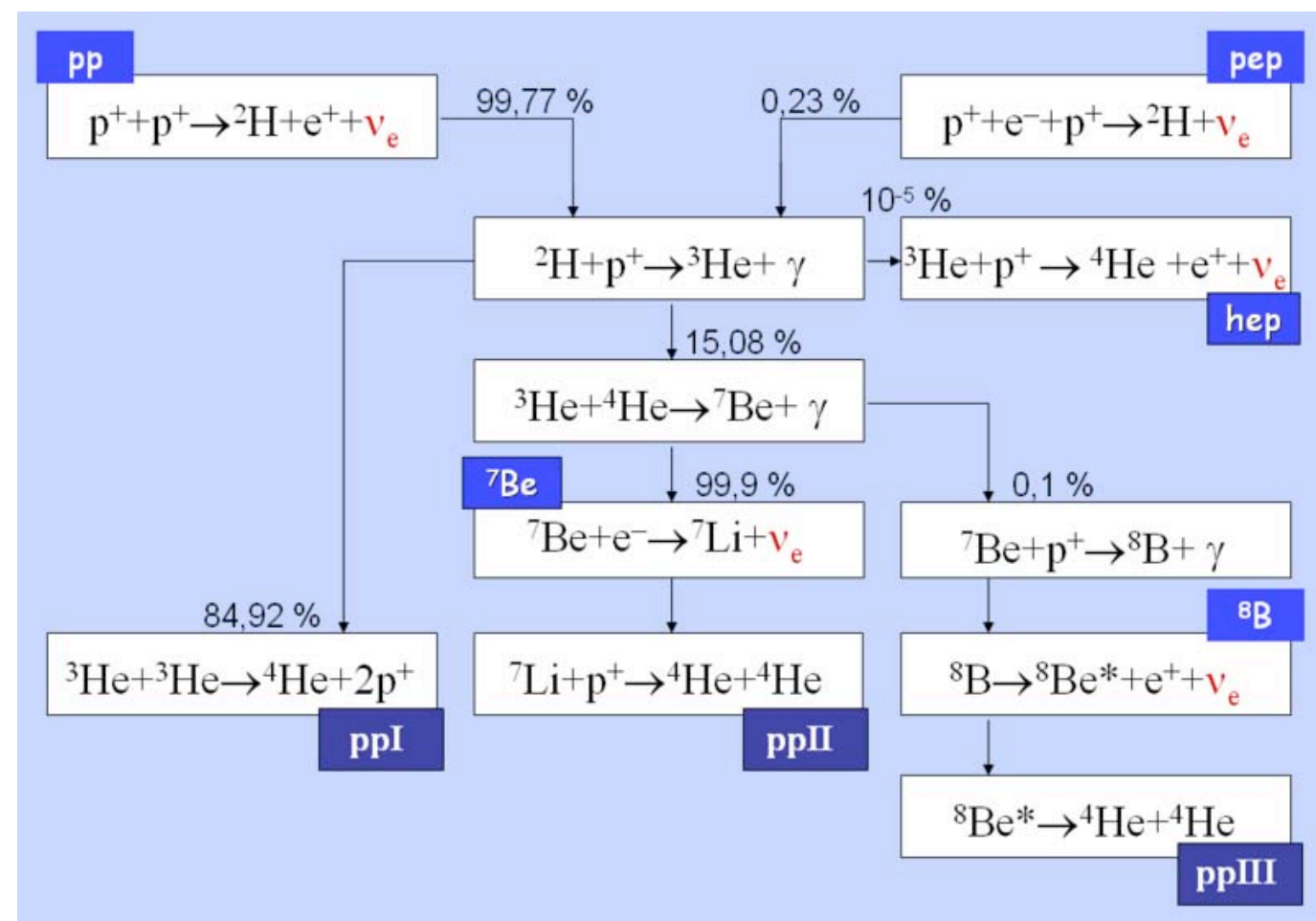
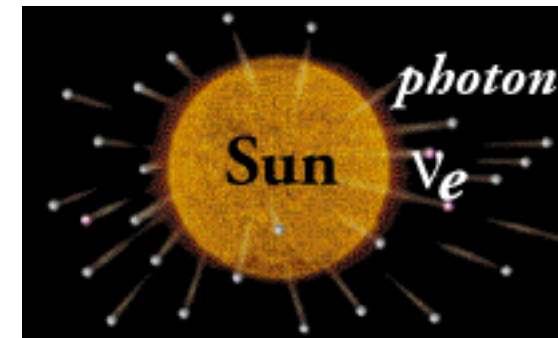


Nuclear binding curve



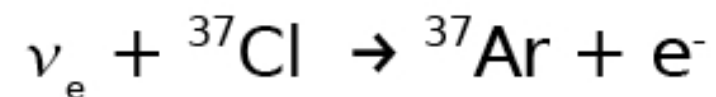
Nuclear reactions: Do you know that...?

- It takes thousands of years for light generated in the solar interior to reach the solar surface (and then another 8 minutes to reach us).
- The Sun generates energy in its core by fusing protons into ^4He through (mostly) the ***pp chain***.
- To prove that the Sun and other main sequence stars are powered by fusion, one can search for the neutrinos (which, unlike photons, don't get trapped in the solar interior)....right?



Nuclear reactions: detecting solar neutrinos

- First detection of solar neutrinos by Ray Davis, Jr. was a heroic effort.
- Ray Davis used this tank of cleaning fluid (615 t) C_2Cl_4 to detect solar neutrinos



- Flushed tank once a month, collected a few (if lucky) ${}^{37}\text{Ar}$ atoms in a glass vial (proportional counter) each time, look for its decay back to ${}^{37}\text{Cl}$...for three decades!

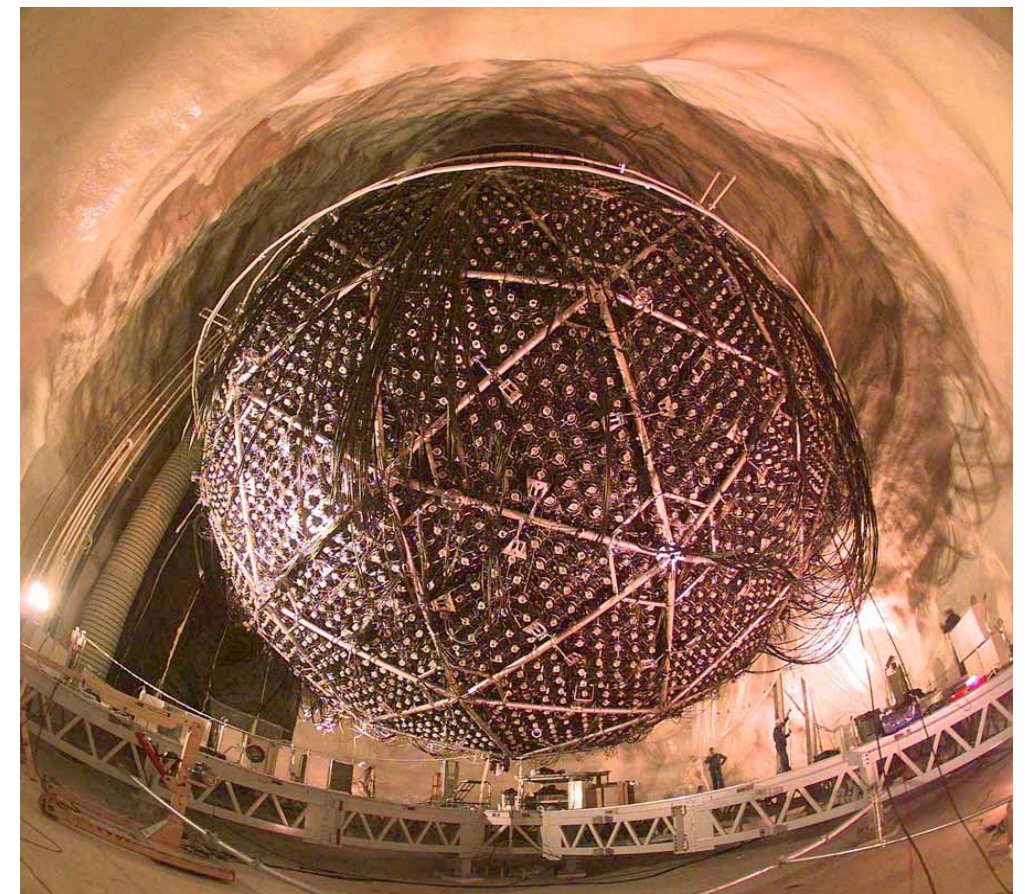


“Ray Davis tells me that the experiment is simple (‘only plumbing’) and that the chemistry is ‘standard.’ I suppose I must believe him, but as a nonchemist I am awed by the magnitude of his task and the accuracy with which he can accomplish it. The total number of atoms in the big tank is about 10^{30} . He is able to find and extract from the tank the few dozen atoms of ${}^{37}\text{Ar}$ that may be produced inside by the capture of solar neutrinos. This makes looking for a needle in a haystack seem easy.” - J. Bahcall

Aside: other neutrino experiments

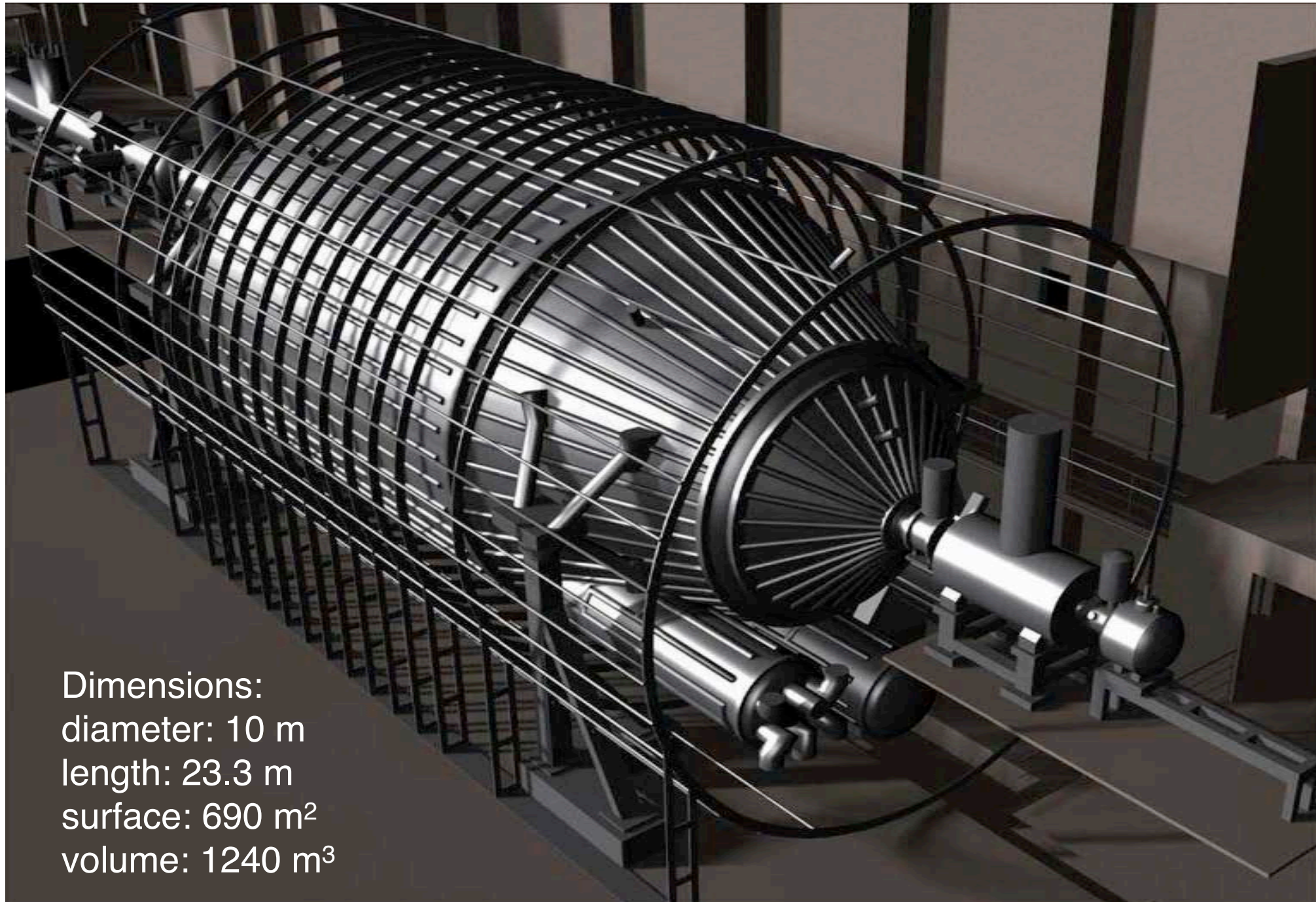


- **Sudbury Neutrino Observatory:** building a cleanroom in an active nickel mine 2 km underground



Aside: other neutrino experiments

- **KATRIN:** the largest ultra-high vacuum tank in the world



Dimensions:
diameter: 10 m
length: 23.3 m
surface: 690 m²
volume: 1240 m³

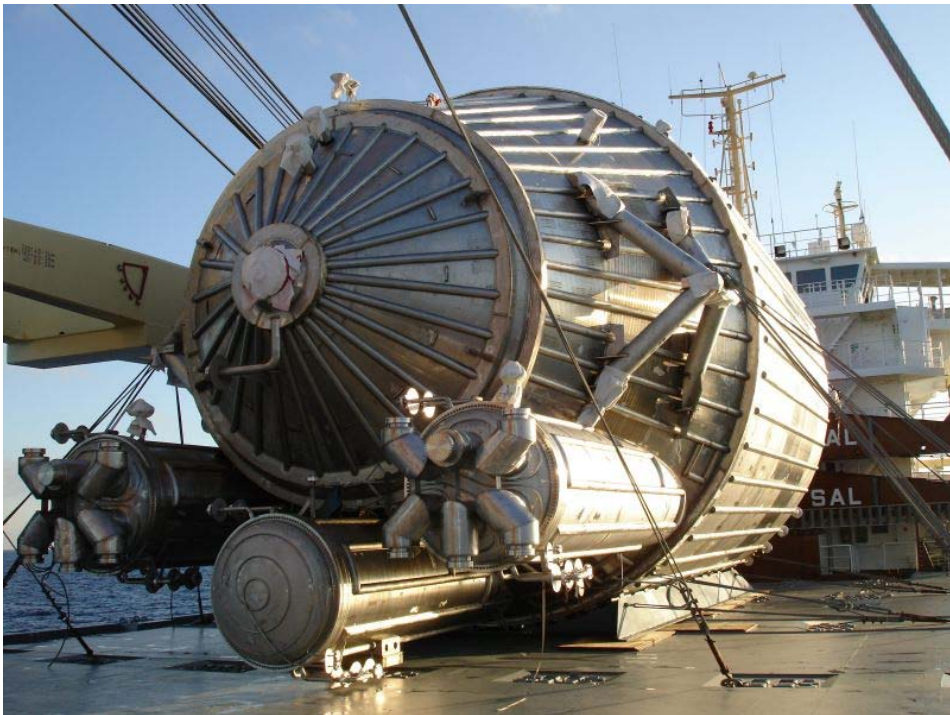
KATRIN



KATRIN



KATRIN



KATRIN

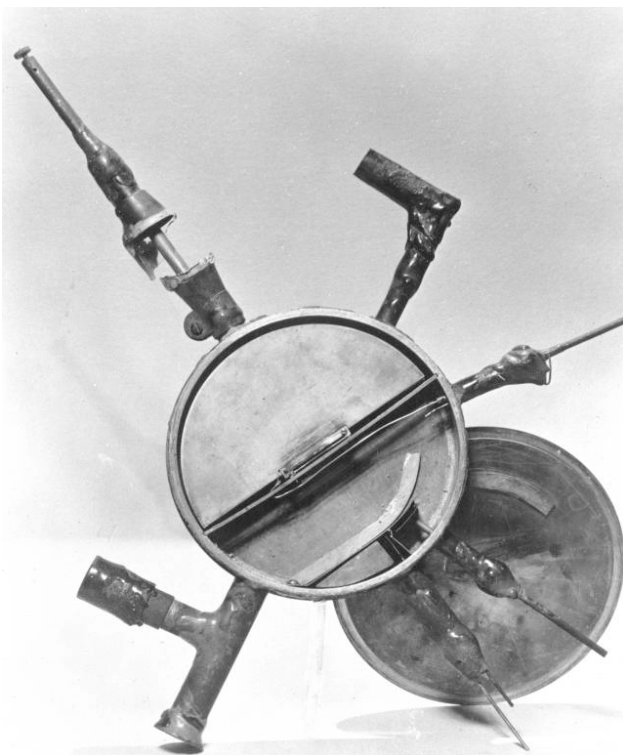


KATRIN

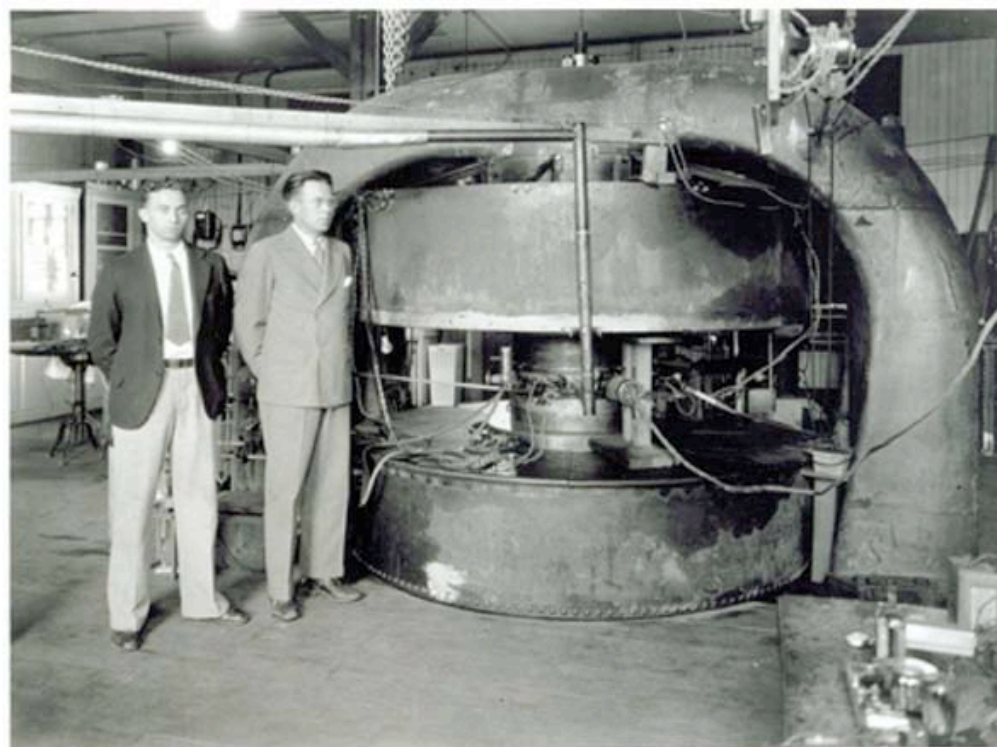


Accelerators: machines for nuclear studies (and a bit of local history)

- Ernest Orlando Lawrence invented the first cyclotron in 1929-31 in a small laboratory on the Berkeley campus. This was the foundation of the “Radiation Laboratory” (and later, Lawrence Berkeley National Lab, my current employer).



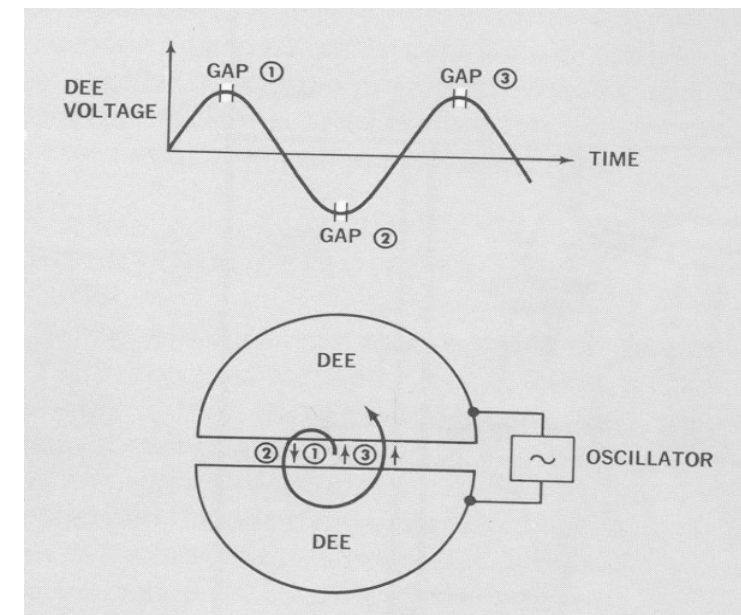
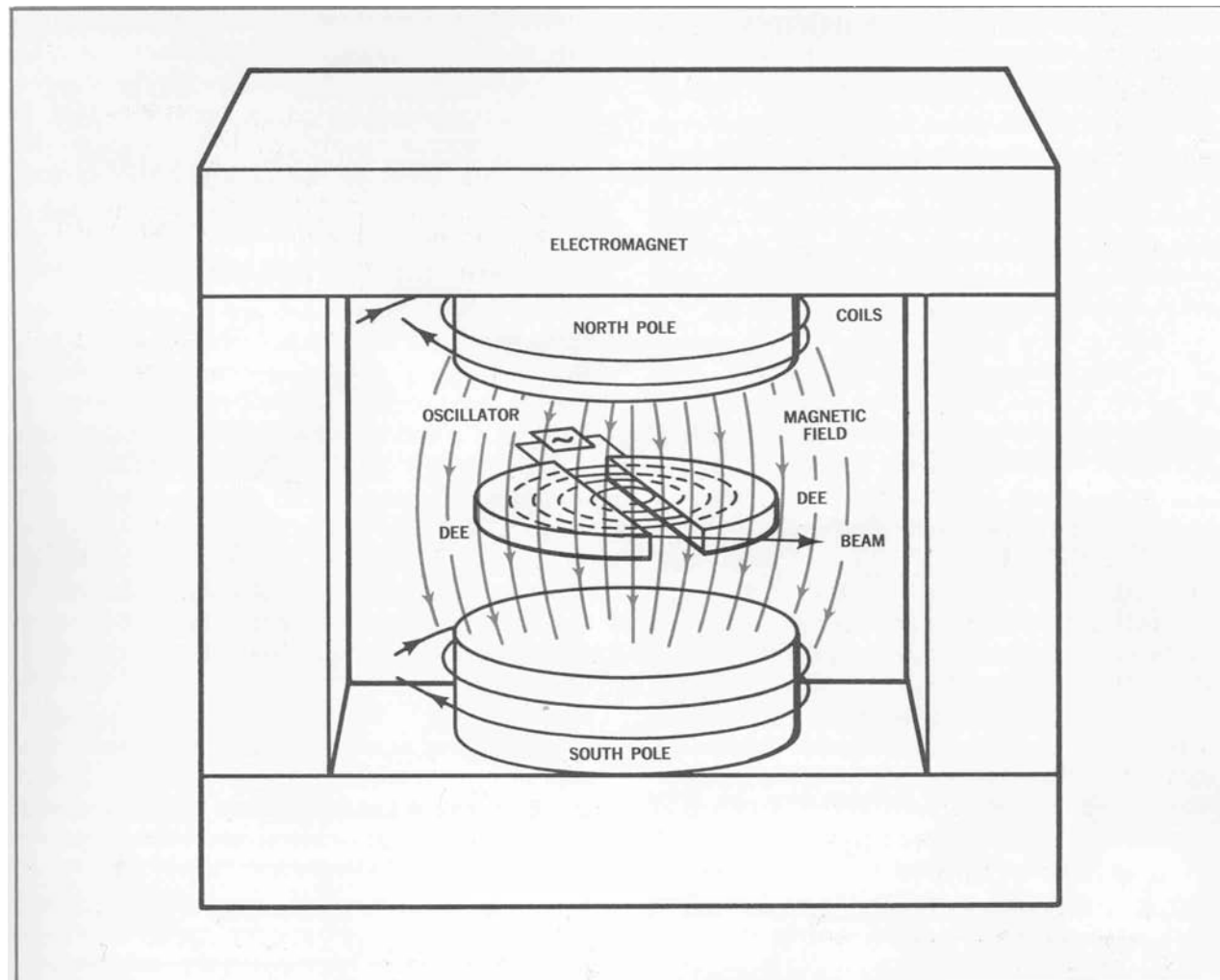
First cyclotron



27" cyclotron; its magnet now outside LHS

- Cyclotrons are still used in physics research, and in cancer therapy and medical isotope production.

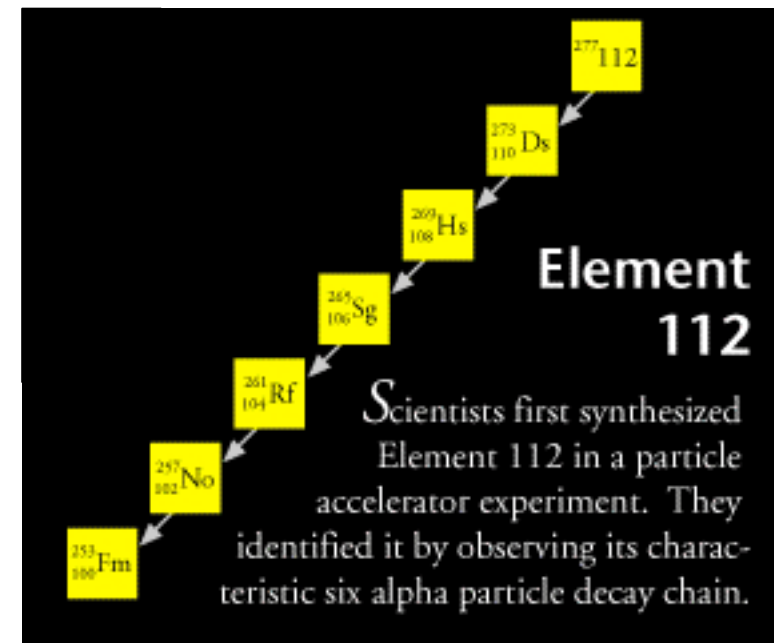
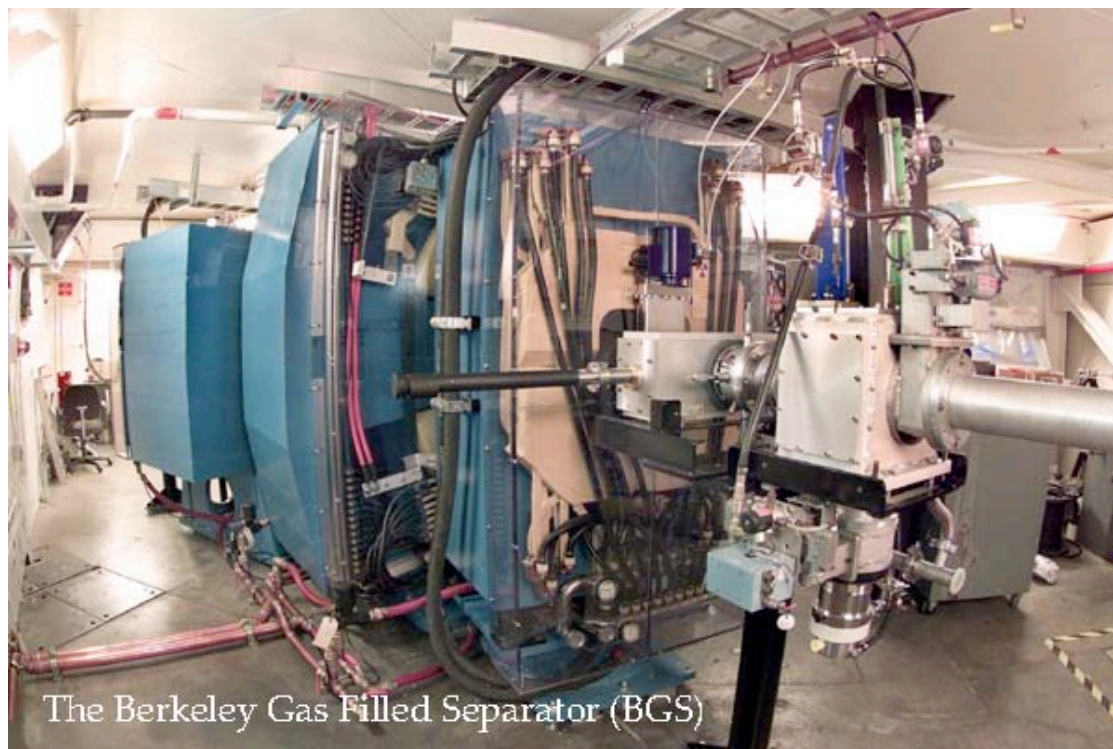
Accelerators: how does a cyclotron work?



- It is left as an exercise for the students to show that the “cyclotron frequency” does not depend on the particles’ orbit radius if the particles are moving much slower than the speed of light.

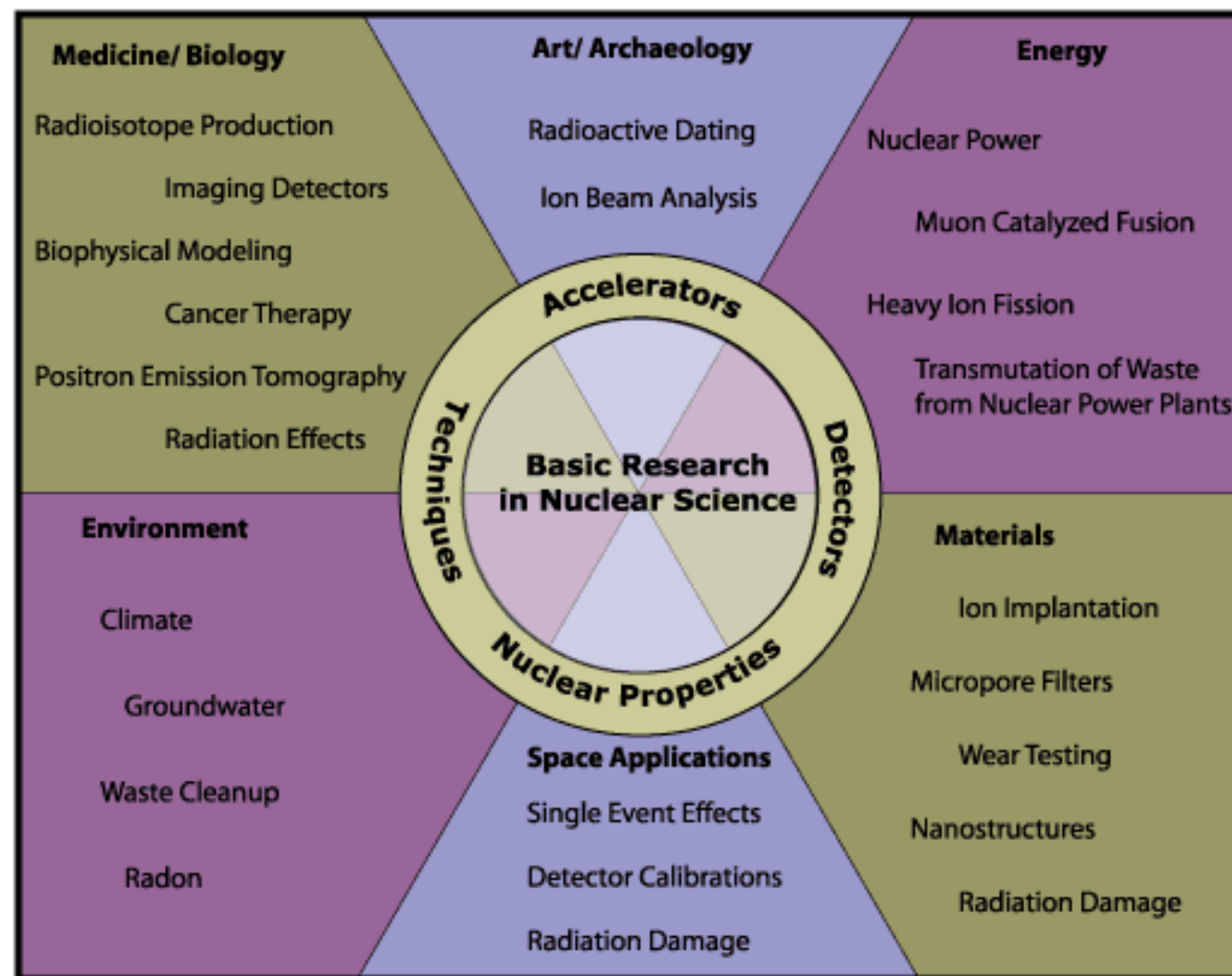
Nuclear reactions: super heavy elements

- Super heavy elements can be “made” via nuclear reaction.
- For example for Z=112: $^{70}\text{Zn} + ^{208}\text{Pb} \rightarrow ^{278}112^*$
- To confirm their existence, first find and catch them, then look for tell-tale sequence of alpha decays.

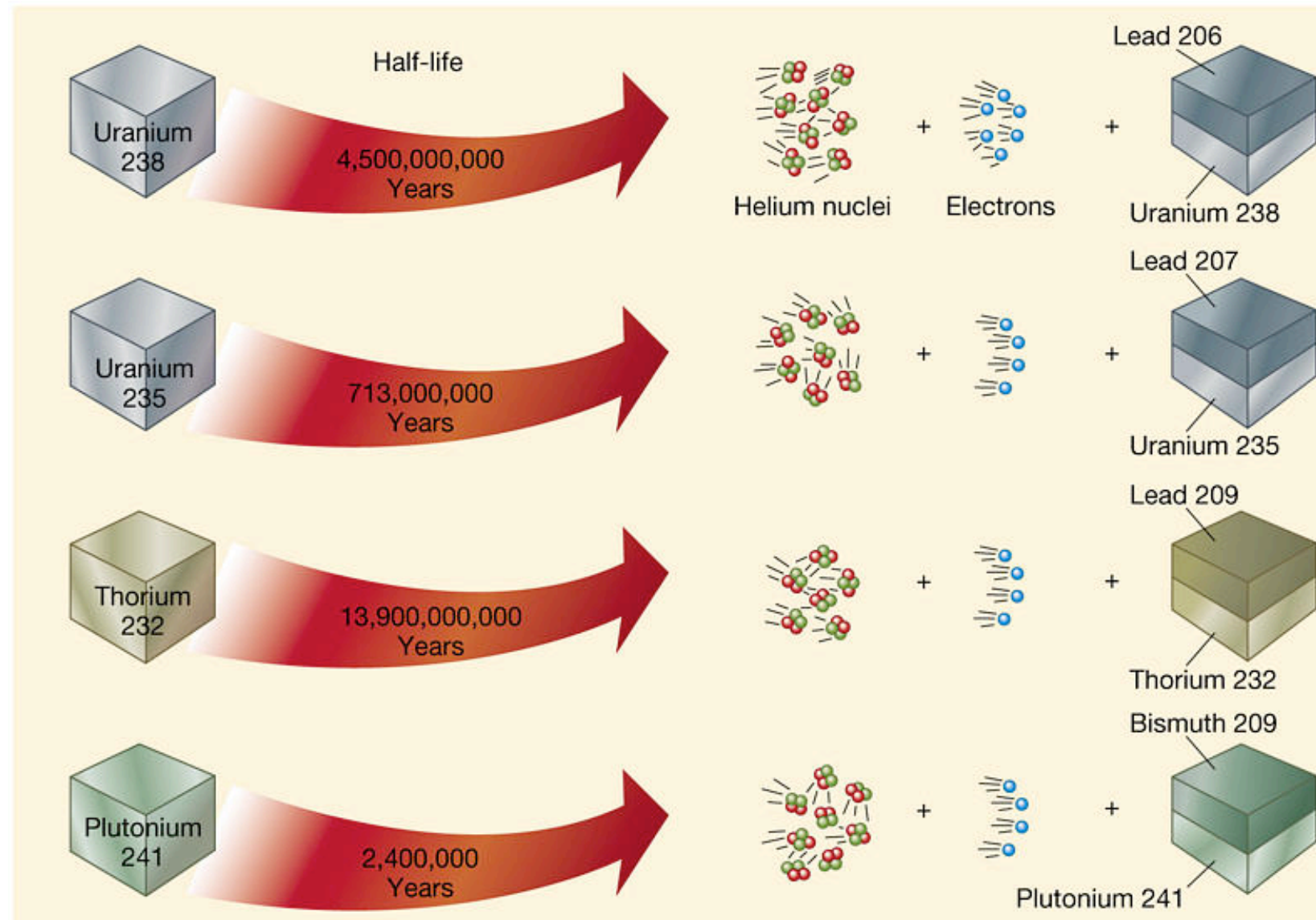


- Element 112 is recently named as Copernicium in honor of Nikolaus Copernicus.

Applications of Nuclear Science



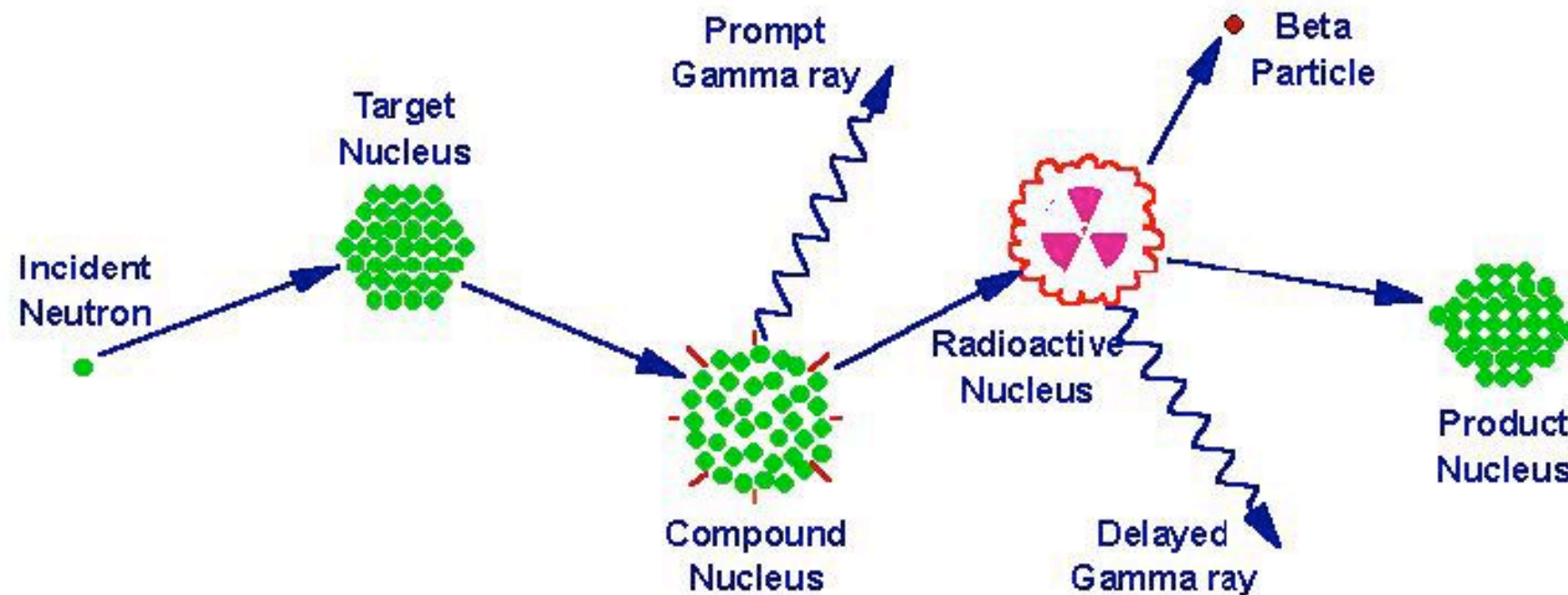
Radioactive dating



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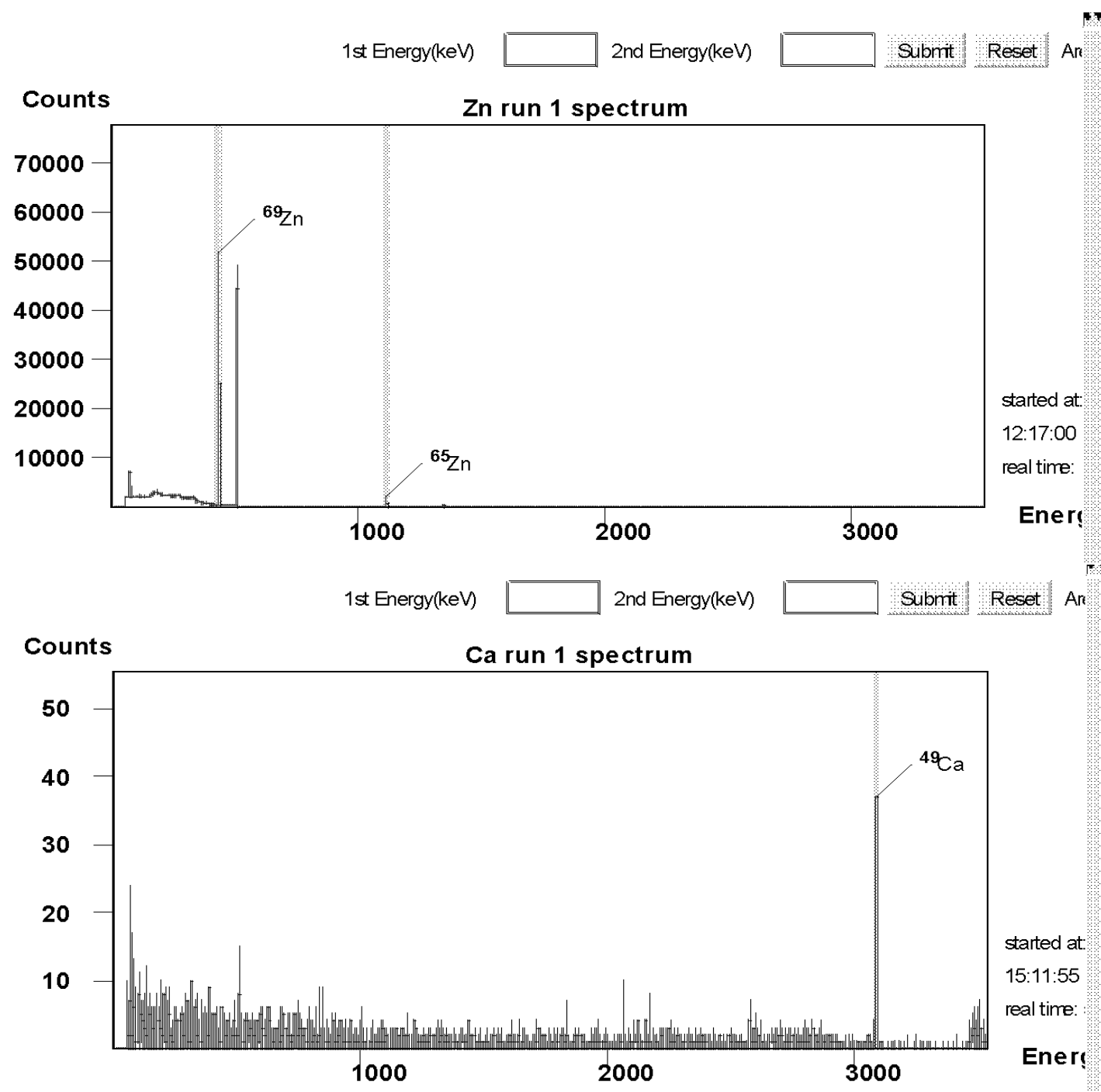
- Use highly sensitive mass spectrometers (modern version of Aston's) to determine parent and daughter nuclei concentrations

Neutron activation analysis



<http://ie.lbl.gov/naa>

Examples of NAA data



Common use of radioactive isotopes

- **Self-luminous “EXIT” signs.** Tritium (^3H) gas is contained in sealed glass tubes, whose interior are lined with a phosphor. Low-energy β emitted by the tritium bombard the phosphor, causing it to glow.
- **Smoke detectors.** The current of alphas (^4He) from ^{241}Am decays entering an ionization chamber is reduced by the presence of smoke particulates, and the alarm is triggered.
- **Radiation sterilization** with ^{60}Co , UV light or X-rays. The radiation causes severe damage to the cell chromosomes, specifically the DNA, in bacteria.



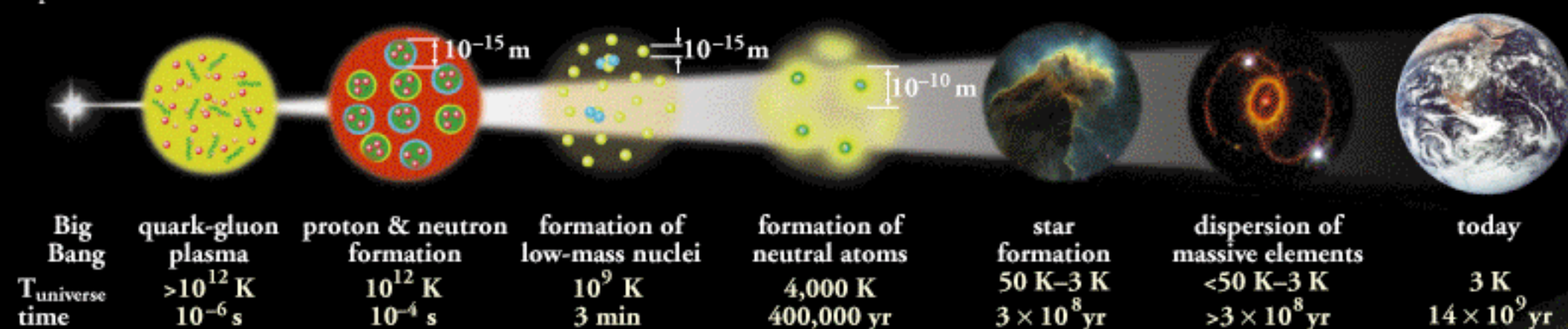
IMAGE STOLEN FROM THEEXITSTORE.COM



Our Universe

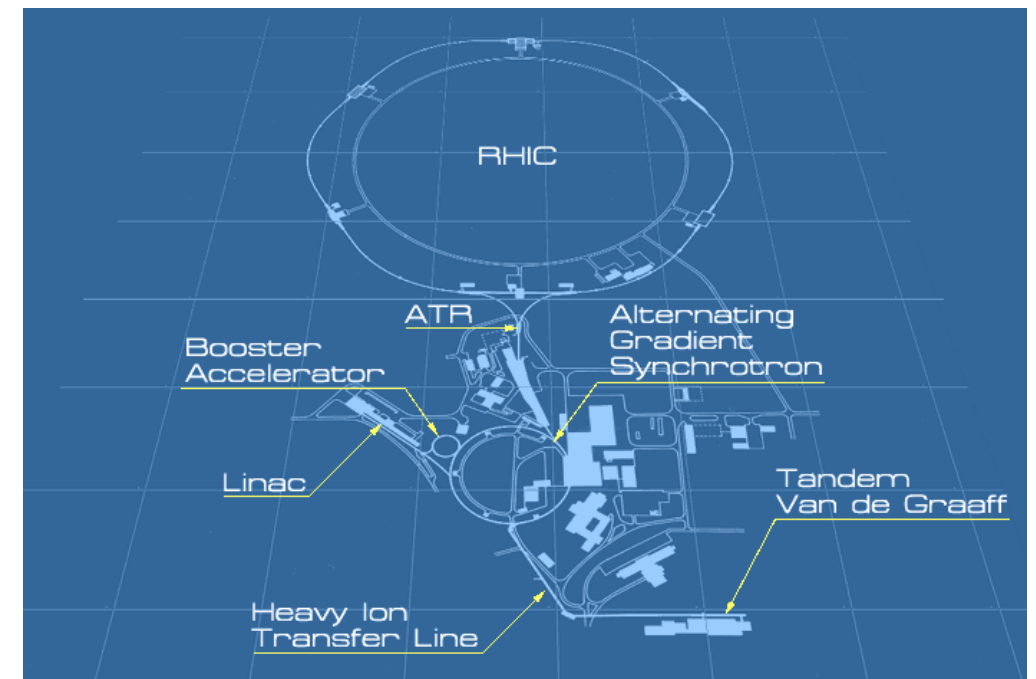
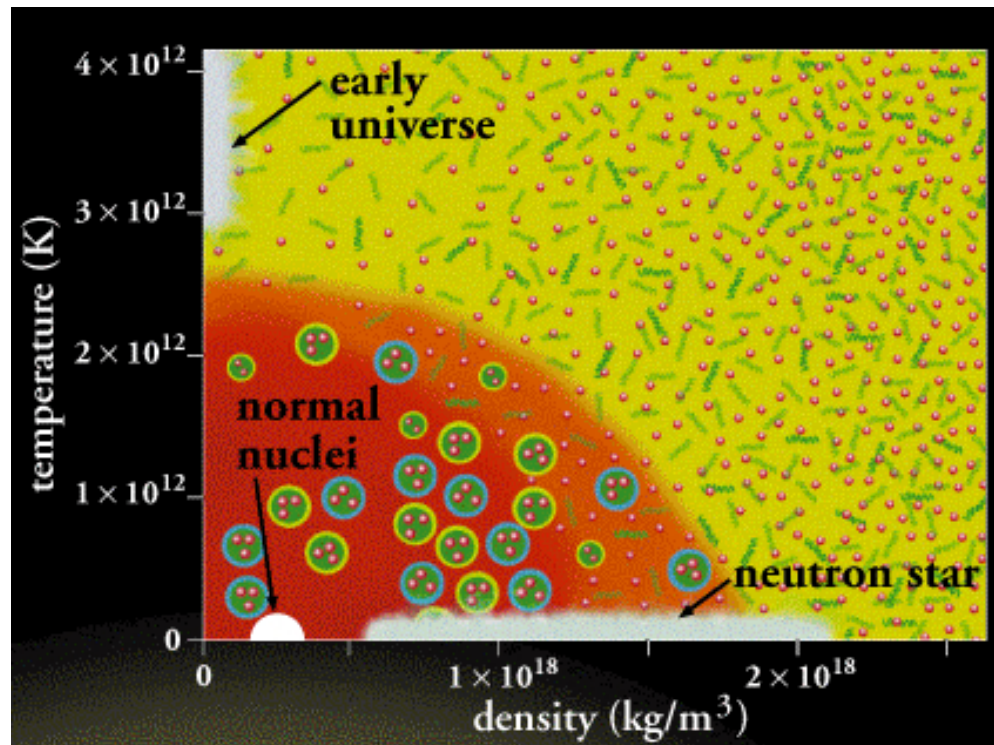
Expansion of the Universe

After the Big Bang, the universe expanded and cooled. At about 10^{-6} second, the universe consisted of a soup of quarks, gluons, electrons, and neutrinos. When the temperature of the Universe, T_{universe} , cooled to about 10^{12} K, this soup coalesced into protons, neutrons, and electrons. As time progressed, some of the protons and neutrons formed deuterium, helium, and lithium nuclei. Still later, electrons combined with protons and these low-mass nuclei to form neutral atoms. Due to gravity, clouds of atoms contracted into stars, where hydrogen and helium fused into more massive chemical elements. Exploding stars (supernovae) form the most massive elements and disperse them into space. Our earth was formed from supernova debris.

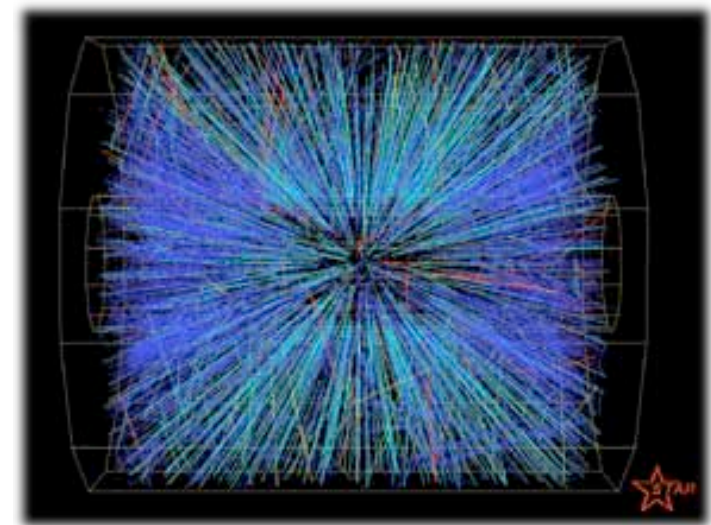


Quark-gluon plasma (QGP)

- **QGP:** A soup of (almost) free quarks and gluons that exists at extremely high temperature and/or density.



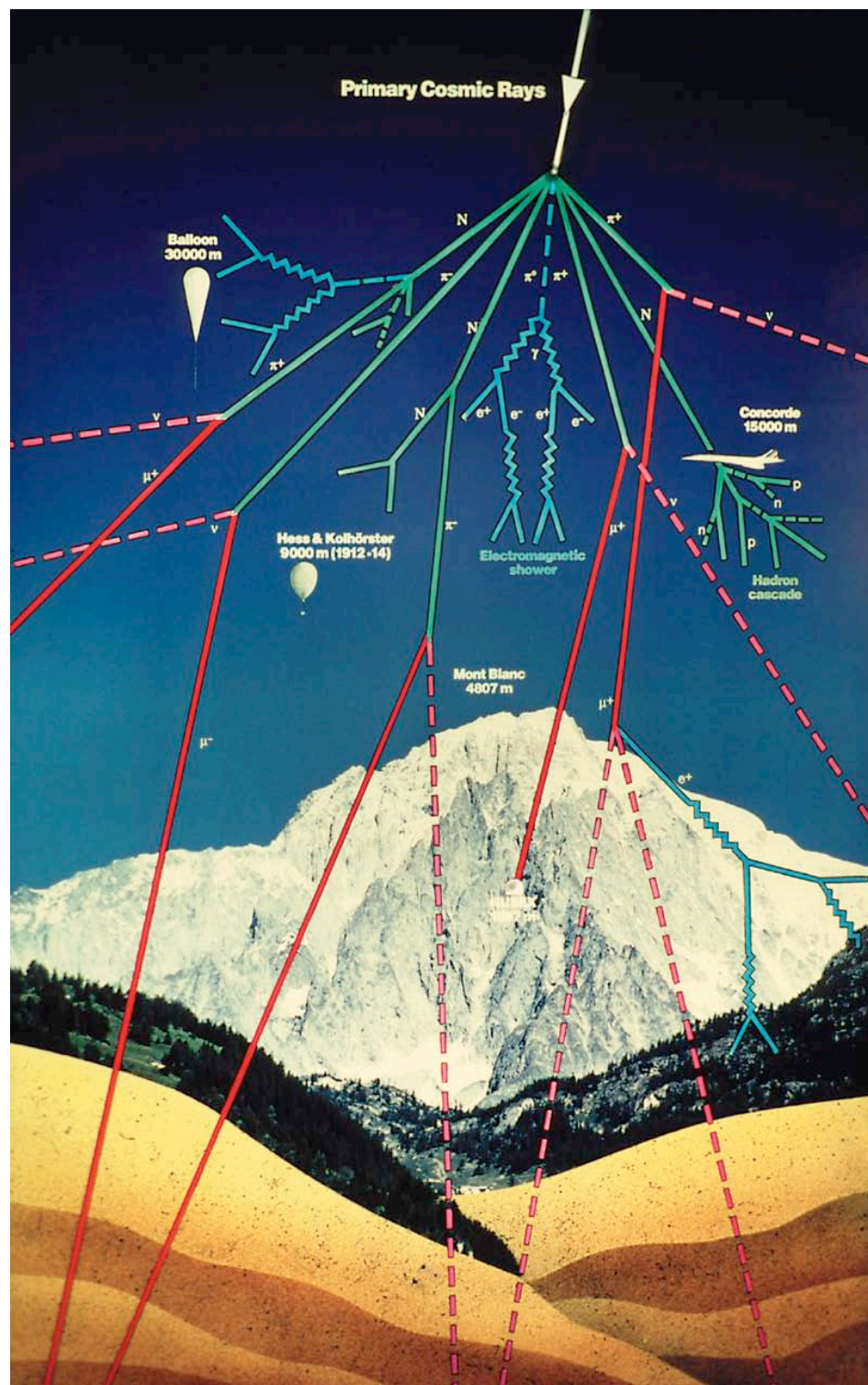
RHIC collider at Brookhaven National Lab



Au + Au collision in STAR detector at RHIC

More details in: <http://www.youtube.com/watch?v=kXy5EvYu3fw>

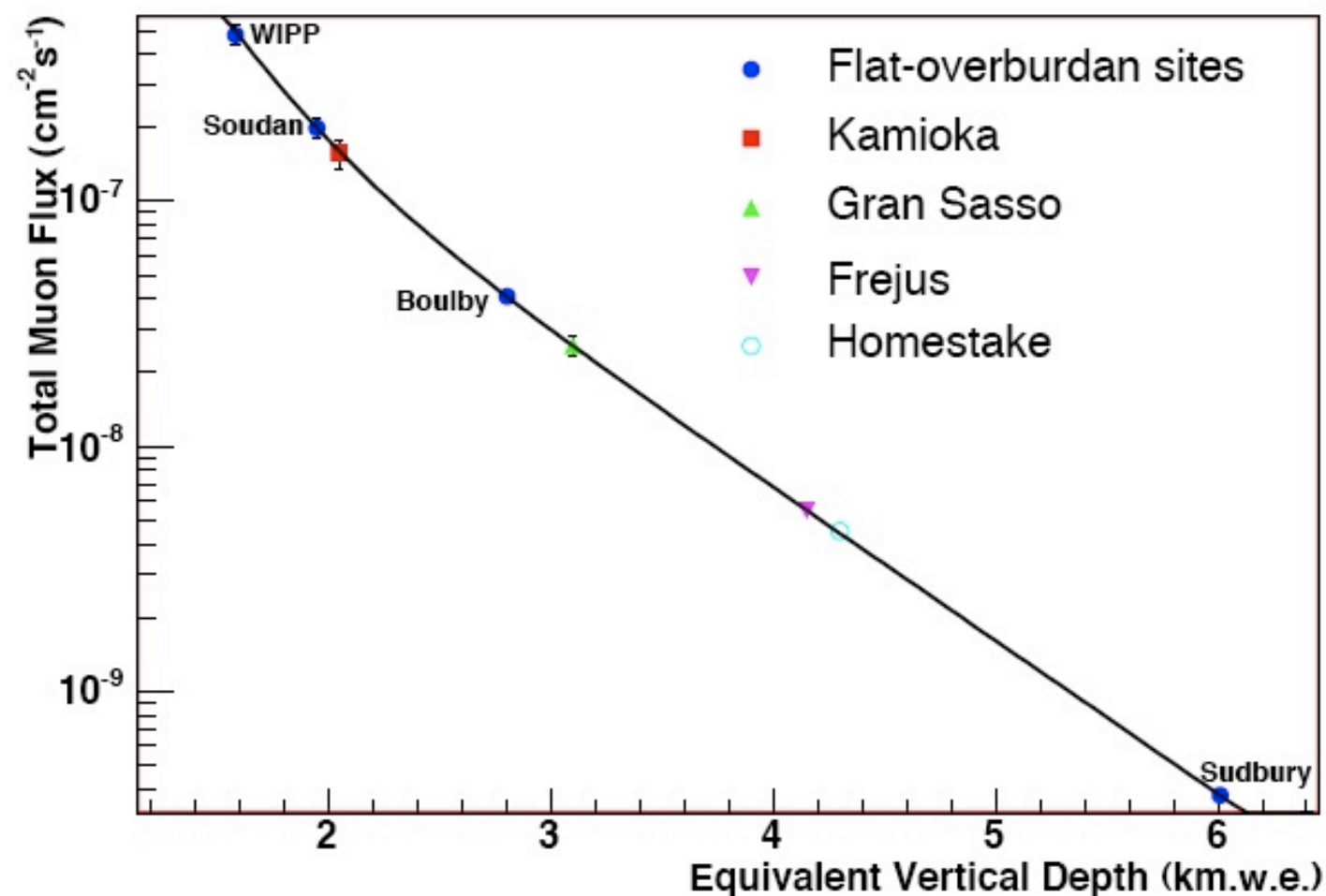
Messengers from the cosmos...



- High energy particles (mostly protons) from galactic and extra-galactic sources bombard our Earth's upper atmosphere.
- They interact with nuclei in the atmosphere (nitrogen, oxygen...) and create a shower of other particles.
- A significant fraction of these secondary particles are muons, a heavier cousin of electrons.

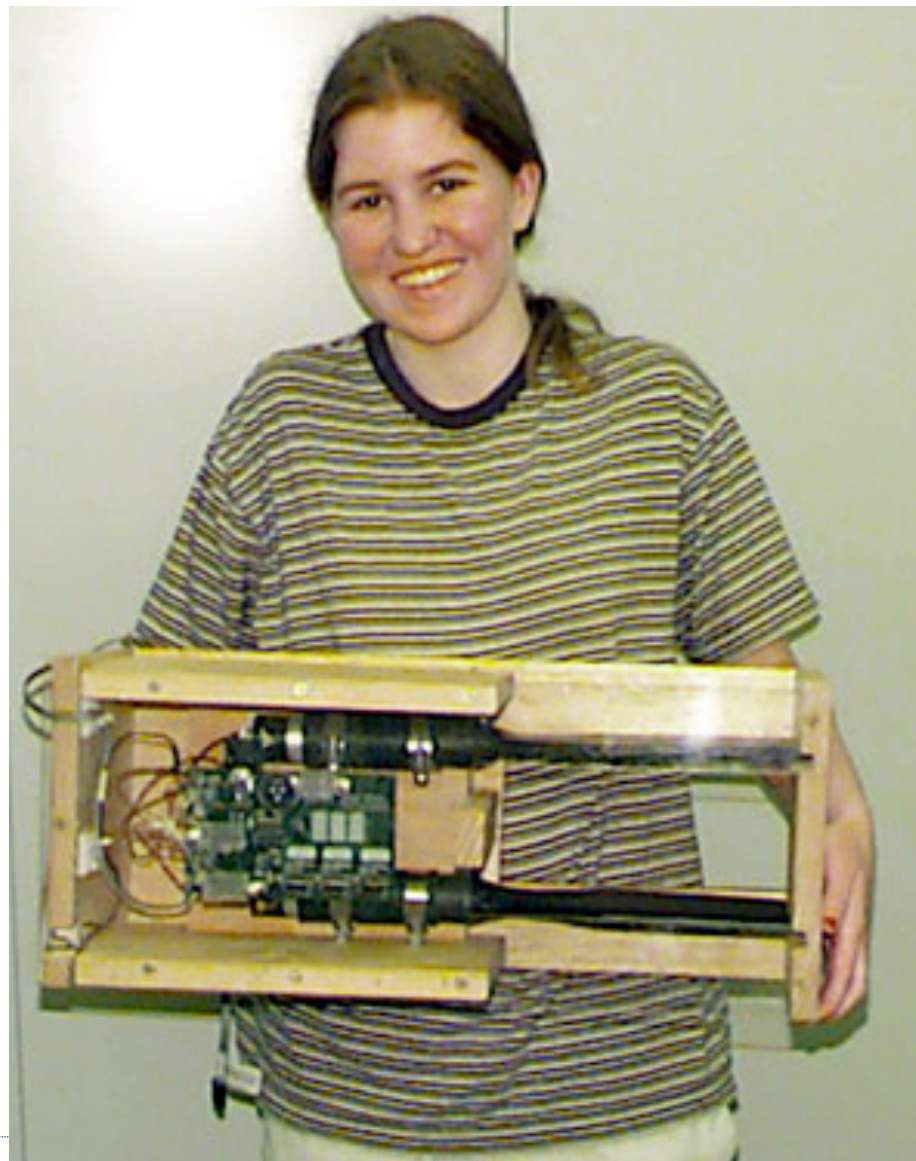
Shielding from cosmic rays?

- Higher elevation → higher exposure to cosmic rays
- To reduce cosmic ray exposure, go underground. In fact, all solar neutrino experiments (remember Ray Davis and his tank of cleaning fluid?) were installed in underground laboratories.



Cosmic rays

or you can take a detector up on a mountain...



The Cosmic Connection

<http://www.lbl.gov/abc/cosmic/>



Thank you!